

Improving Covariance Matrices using Machine Learning

Natalí Soler Matubaro de Santi

natalidesanti@gmail.com

Advisor: Prof. Dr. Luis Raul Weber Abramo

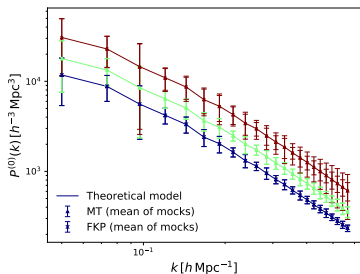
Mathematical Physics Department
University of São Paulo

December 17, 2020



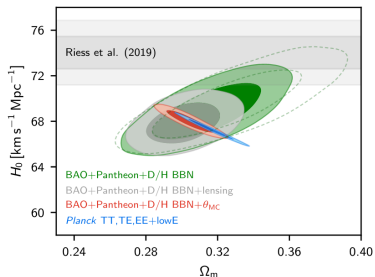
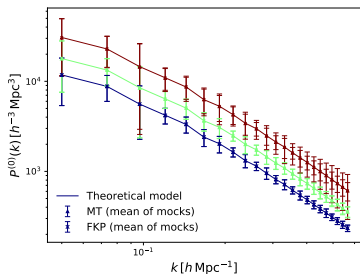
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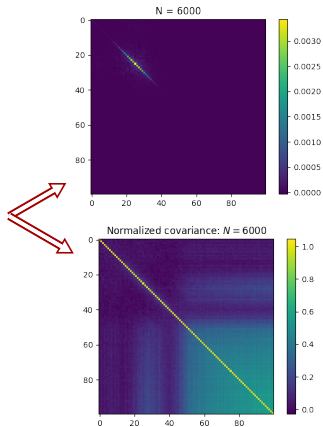
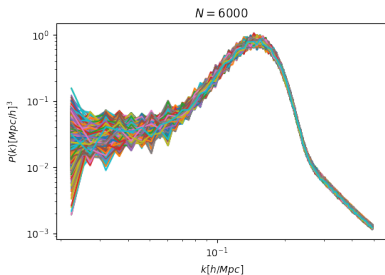
- They reflect the propagation of the statistical errors:



Planck 2018 results: VI Cosmological parameters.

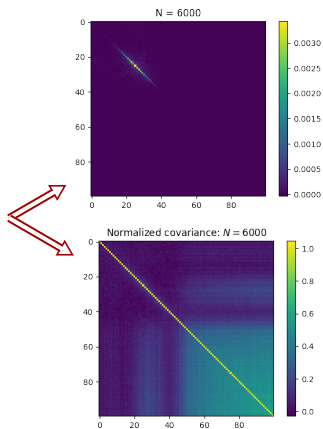
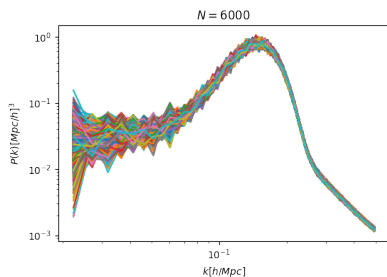
Problem

- To represent the true statistical errors we need a lot of data to build these matrices:



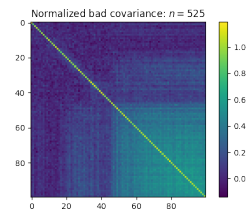
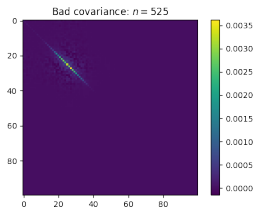
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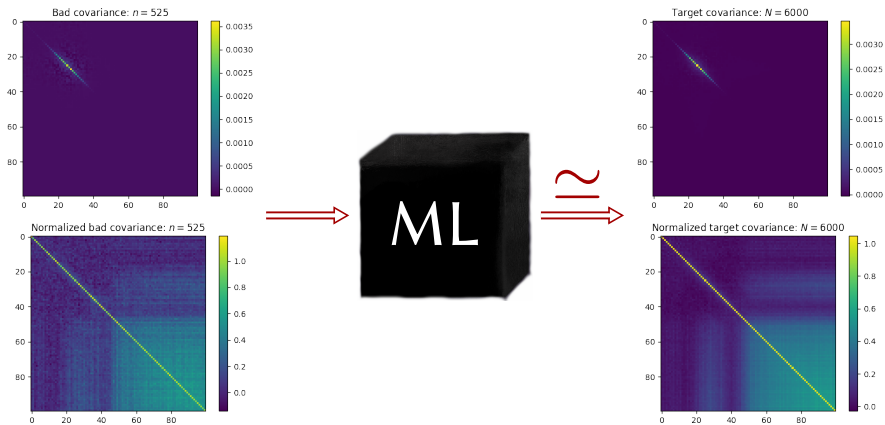


But we can not always do it in practice...

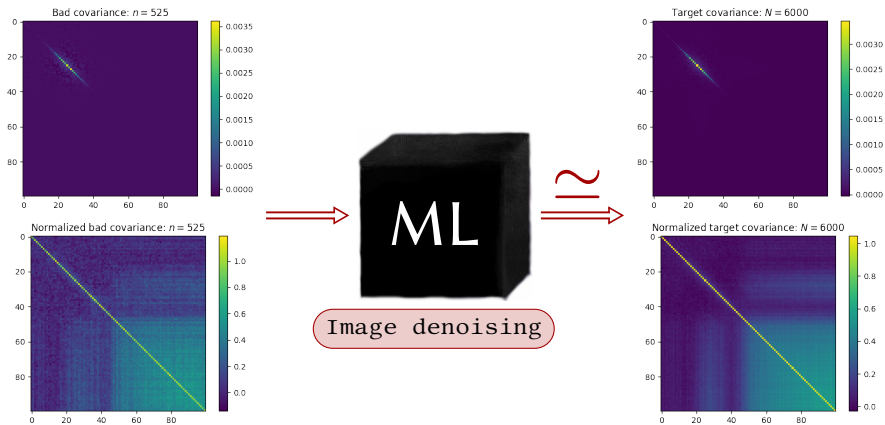
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Methodology - Toy project simulation

- The simulations followed the theoretical function:

$$P(k) = P_0 \exp \left[-\frac{(k - k_0)^2}{2\sigma_0^2} \right]$$

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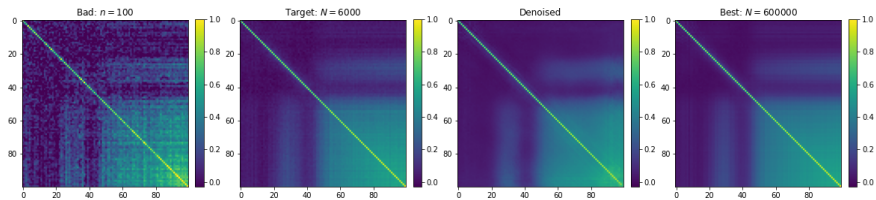
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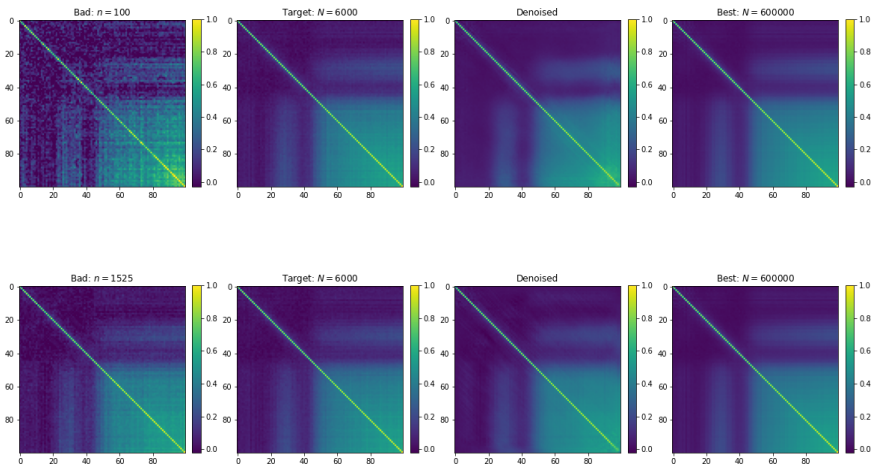
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- Then, we test the each **denoiser** with:
 - `input_test`: **bad cov. matrices** (hundreds of spectra) $\Rightarrow n$.

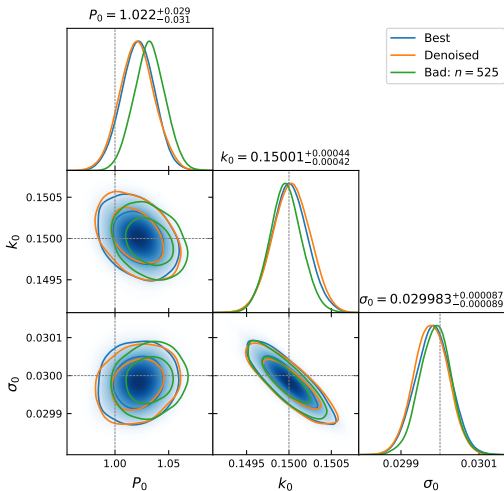
Results - Visual Results



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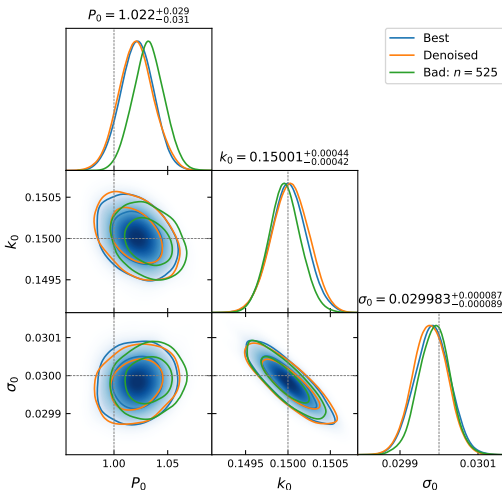


Results - Markov Monte Carlo Chain (MCMC)¹: Recovering Parameters



¹Ivezić, Z. et al., Statistics, Data Mining, and Machine Learning in Astronomy: A Practical Python Guide for the Analysis of Survey Data, 2014.

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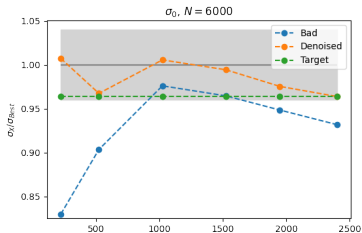
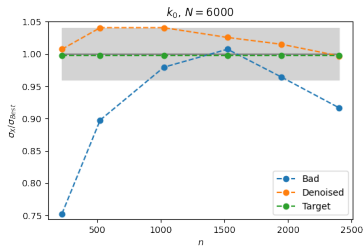
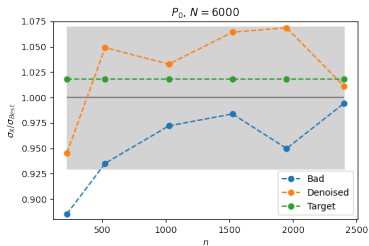
$$\frac{\Delta\mu}{\sigma_{Best}} = \frac{\mu_X - \mu_{Best}}{\sigma_{Best}}$$

$\Delta\mu/\sigma_{Best}$	Bad	Denoised
P_0	0.685	0.054
k_0	0.218	0.092
σ_0	0.118	0.120

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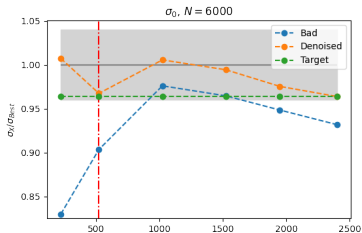
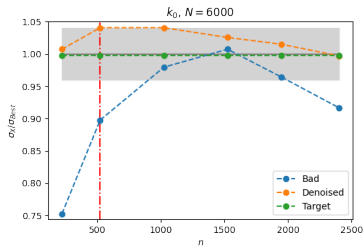
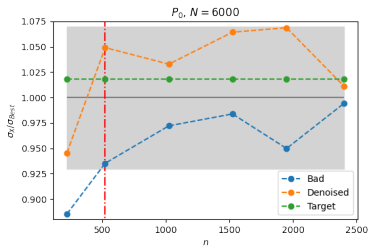
Results - Sigma fraction: σ_X/σ_{Best}

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- We have achieved great results using **image denoising** techniques to improve the *covariance matrices*;
- We showed that even with a **low** number of simulations, we can achieve the same results as a **higher** number of them;
- Once all this work is a really controlled “toy project”, we want to apply the same method in **realistic simulations** (e.g., ExSHalos, LogNormals, N-body).