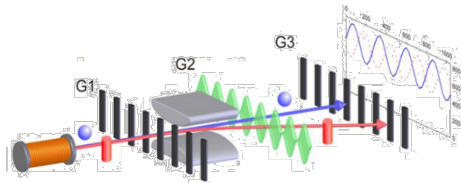
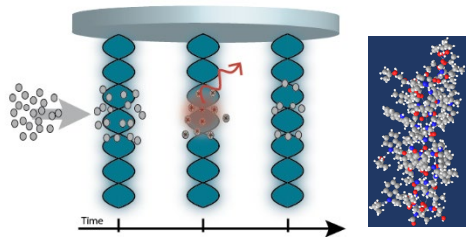


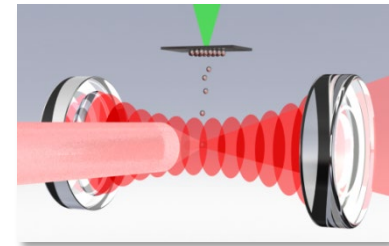
Universal Matter-Wave Interferometry & Gravitational Physics



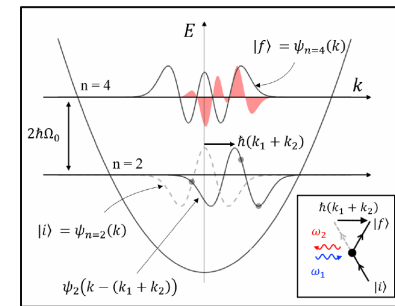
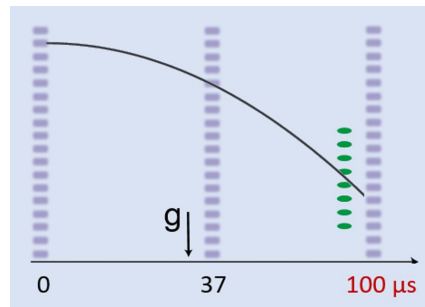
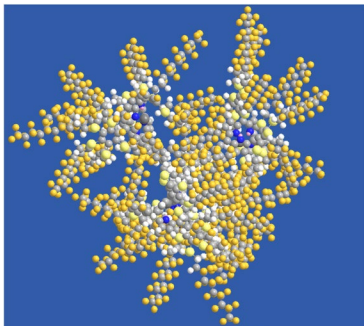
LUMI-Interferometry



OTIMA-Interferometry



Trapped Particles



Markus Arndt

University of Vienna, Quantum Nanophysics Group

www.quantumnano.at



European Research Council
Established by the European Commission





Far-field:

- Ksenjia Simonović

OTIMA

- Philip Rieser
- Armin Shayeghi

CAVITY COOLING

- Stephan Troyer
- Nafia Rahaman
- Stefan Putz

THEORY

- Filip Kialka
- Aljosha Vukovic

Bio-beams/SNWDs

- Marcel Strauß
- Martin Mauser
- Julia Salapa
- Armin Shayeghi
- Philipp Geyer

LUMI

- Yaakov Fein
- Philipp Geyer
- Sebastian Pedalino
- Stefan Gerlich

Former Members

Far-field:

- Thomas Juffmann
- Christian Brand
- Adriana Milic

KDTLI

- Lucia Hackermüller
- Joseph Cotter
- Lukas Mairhofer
- Hendrik Ulbricht
- Sandra Eibenberger

OTIMA

- Philipp Haslinger
- Nadine Dörre
- Jonas Rodewald

Former Members

CAVITY

- Peter Asenbaum
- Stefan Kuhn
- James Millen
- Pietro Vahramian

Bio Sources

- Maxime Debiossac
- Moritz Kriegleder
- Ugur Sezer

THEORY

- Stefan Nimmrichter

Collaborations

Duisburg: Theory

Klaus Hornberger, Benjamin Stickler

Basel: Chemistry

Marcel Mayor, Valentin Köhler

Tel Aviv: Nanofabrication & Imaging

Ori Cheshnovsky, Fernando Patolsky

Uni Wien: Micro-mirror fabrication

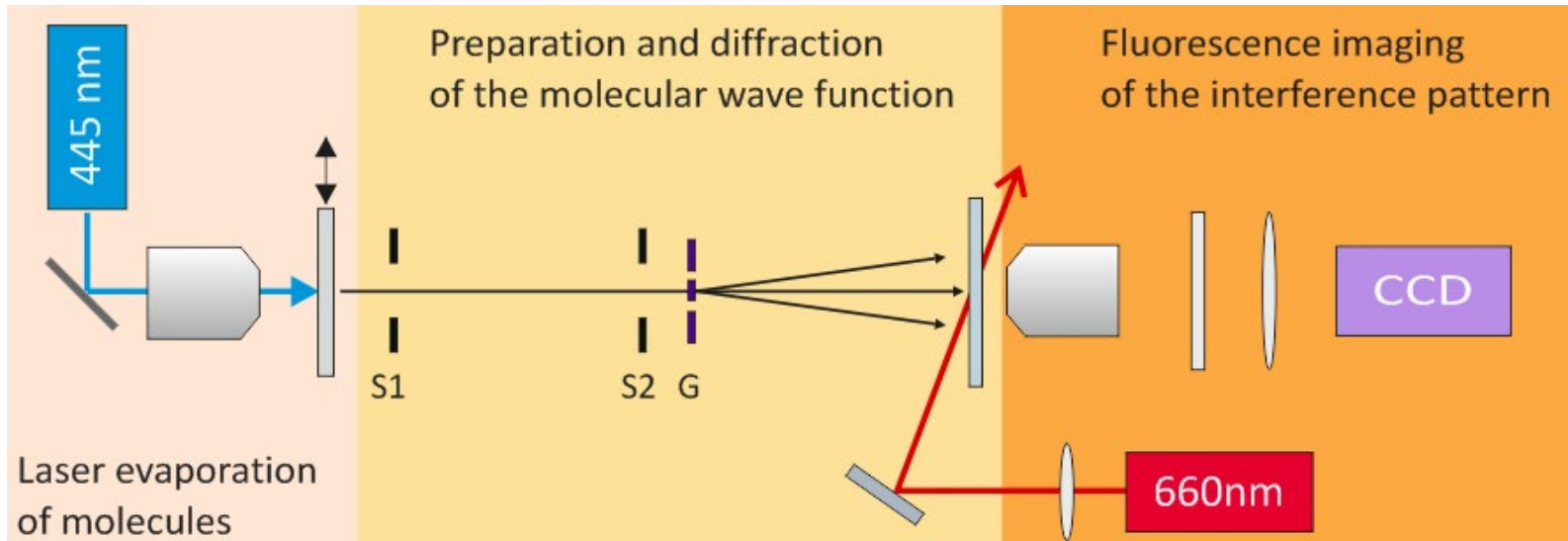
Michael Trupke, Georg Wachter

Kapitza-Dirac Blockade

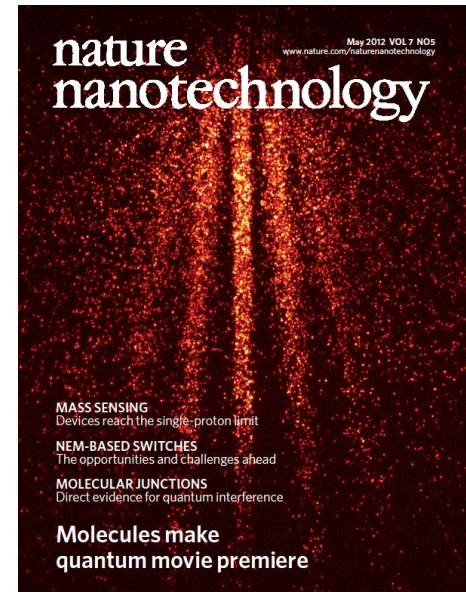
Wayne Cheng-Wei Huang, Göttingen
Herman Batelaan, Lincoln/Nebraska

1. The trivial aspect of gravity in molecule diffraction

Free fall & Gravitational velocity selection



| | |
|--|---|
| | |
| <p>SiN Nanograting</p> <p>$d=100\text{ nm}$</p> <p>$s=50\text{ nm}$</p> <p>$t=10\text{ nm}$</p> | <p>Phthalocyanin</p> <p>$m = 514\text{ amu}$</p> <p>$v = 100\text{ m/s}$</p> <p>$\lambda_{dB} = 5\text{ pm}$</p> |



2. A long-term research programme: Dephasing by stochastic gravitational wave background ?

- Matter-wave interference contrast reduced by **stochastic gravitational wave background**

$$V = \exp\left(\frac{i\Delta\varphi^2}{2}\right) \cdot V_0$$

with an rms phase perturbation

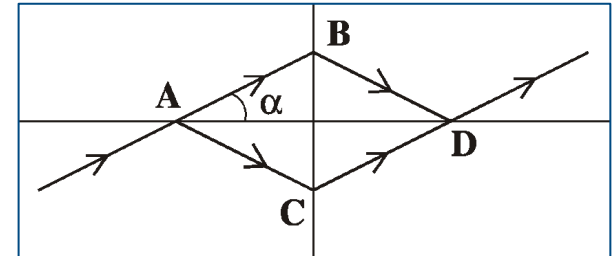
$$\Delta\varphi^2 = \int_0 \frac{d\omega}{2\pi} S_h[\omega] \cdot A[\omega] \cdot F[\omega]$$

where

$$A[\omega] = 16 \cdot \Omega^2 \sin^2(\alpha) \cdot (1 - \cos(\omega\tau))^2 / \omega^2$$

- Quadratic in kinetic energy $\Omega^2 = E_{kin}^2 / \hbar^2$
- Too small for cold-atom & high-mass interferometry (by many orders of magnitude)
- But:** for any **experimentally accessible** λ_{dB} the **lightest mass** serves the purpose best

$$E_{kin} = \frac{p^2}{2m} = \frac{h^2}{2m\lambda_{dB}^2}$$

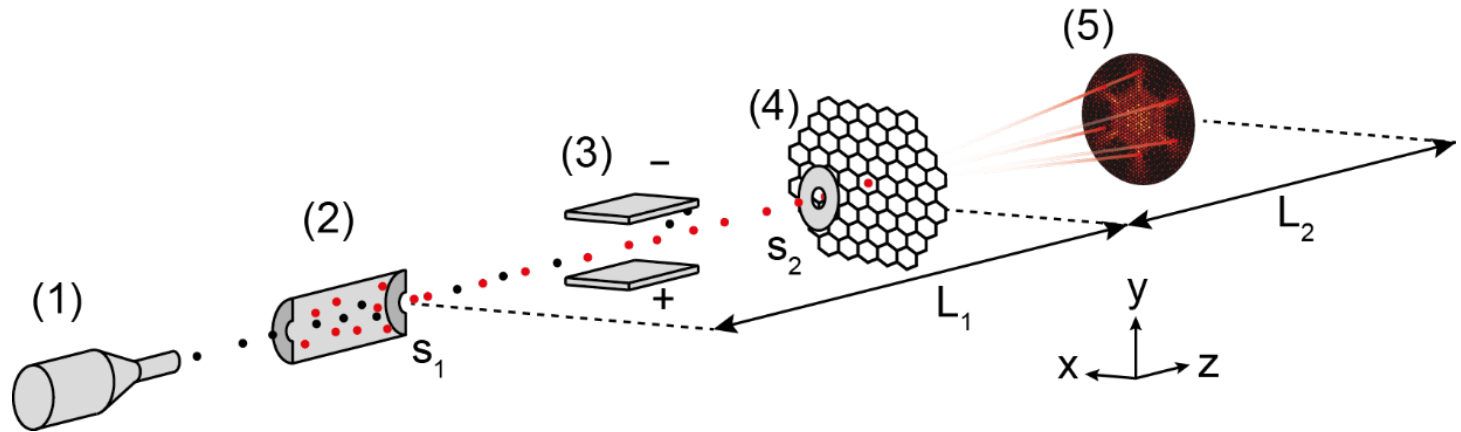


- $S_h[\omega]$ = gravitational strain noise
- $F[\omega]$ = filter function
- $A[\omega]$ = apparatus function

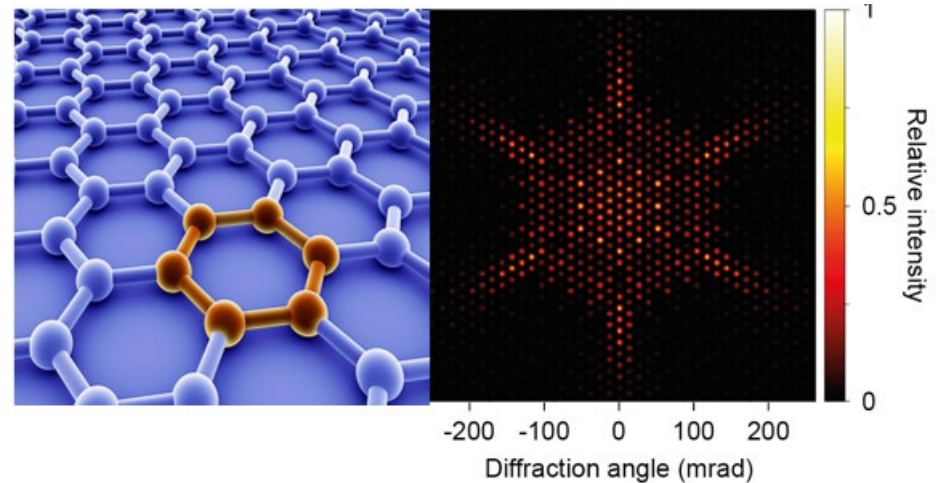
Is it possible to generate wide angle beam splitter for superfast atoms?



Christian Brand,
now DLR Ulm



- Hydrogen atoms
- $E = 80 \text{ eV}$
- $v = 124 \text{ km/s}$
- $\lambda_{dB} = 3.4 \text{ pm}$
- Graphene lattice: $d=2.46 \text{ \AA}$



Currently still a thought experiment ...

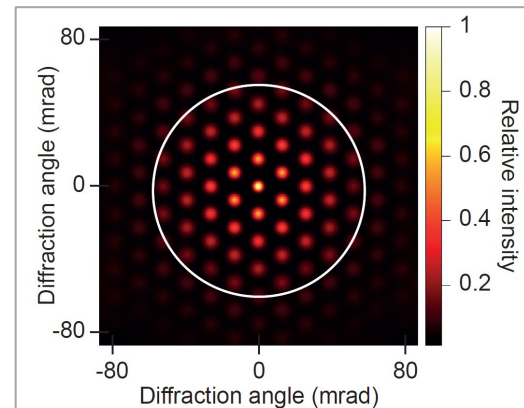
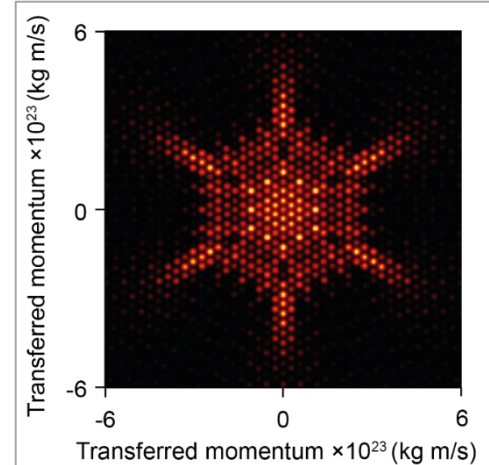
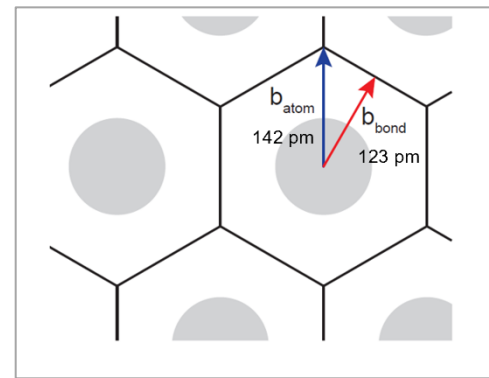
For atoms with $E = 80$ eV transmitted in the grey area

- Δp to carbon mesh < intrinsic momentum uncertainty
- 8-10% of atoms **diffract coherently** up < 50 mrad (4th order)
- In $\tau = 100$ ms: **2700 m** flight distance, **5 cm** falling height
- Coherent beam splitting of **135 m** conceivable (?!!)

Thought experiment:

- $M = 1$ amu, $v = 10^5$ m/s $\rightarrow \lambda = 4$ pm
- $d = 2.4 \times 10^{-10}$ m, $\alpha \simeq 10$ mrad, $L = 4000$ m,
 $\rightarrow E = 80$ eV, $A = 10$ m², $\tau = 10^{-2}$ s

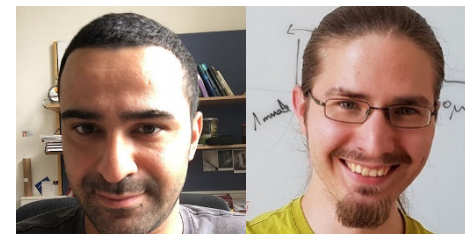
Feasible in the long run ?



**A weak equivalence
demonstrator in
OTIMA interferometry**

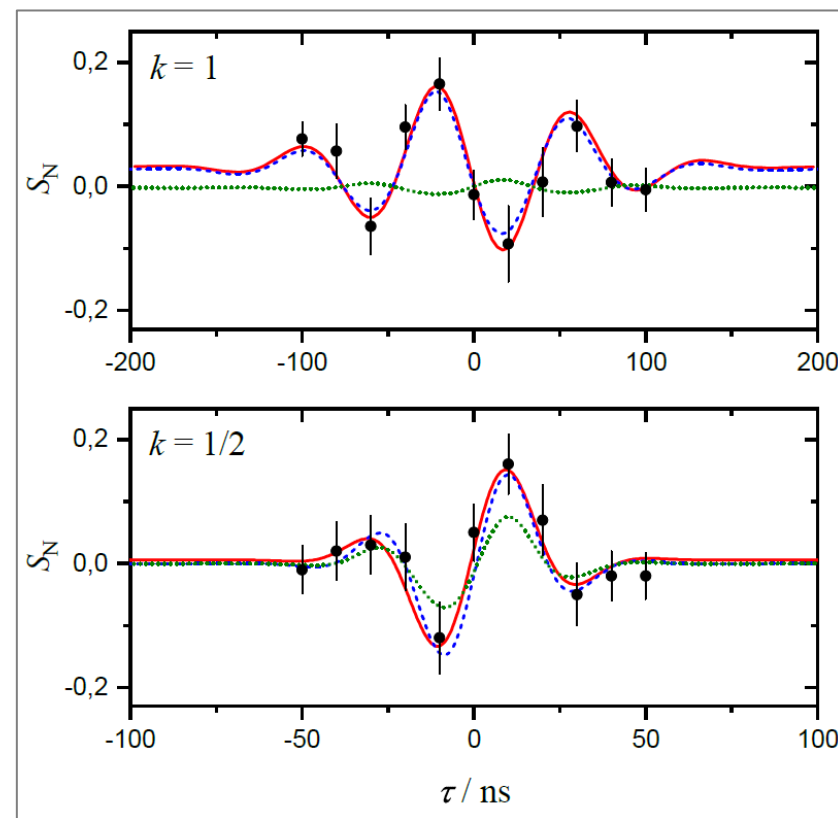
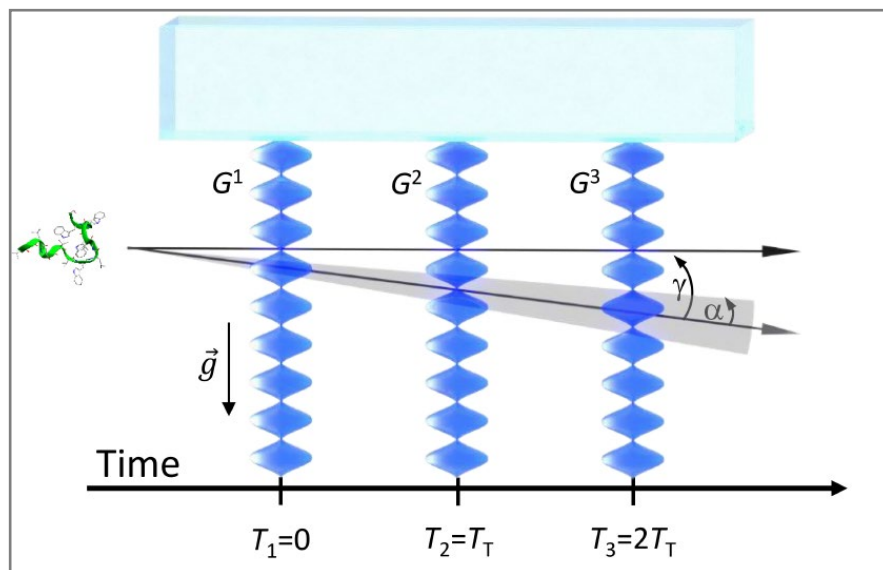
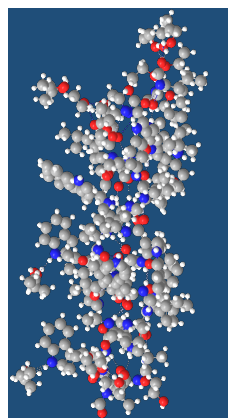
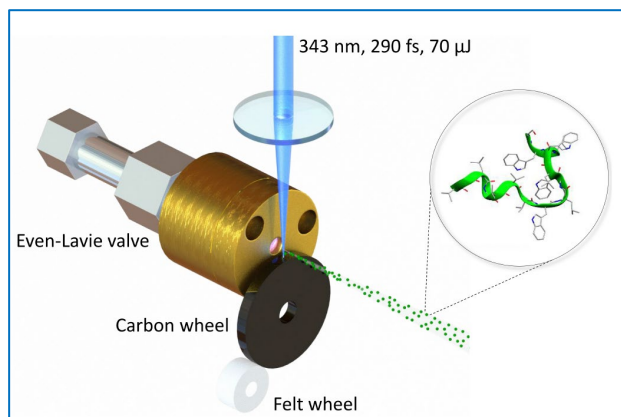
Near field OTIMA interferometry

The quantum wave nature of antibiotic polypeptides

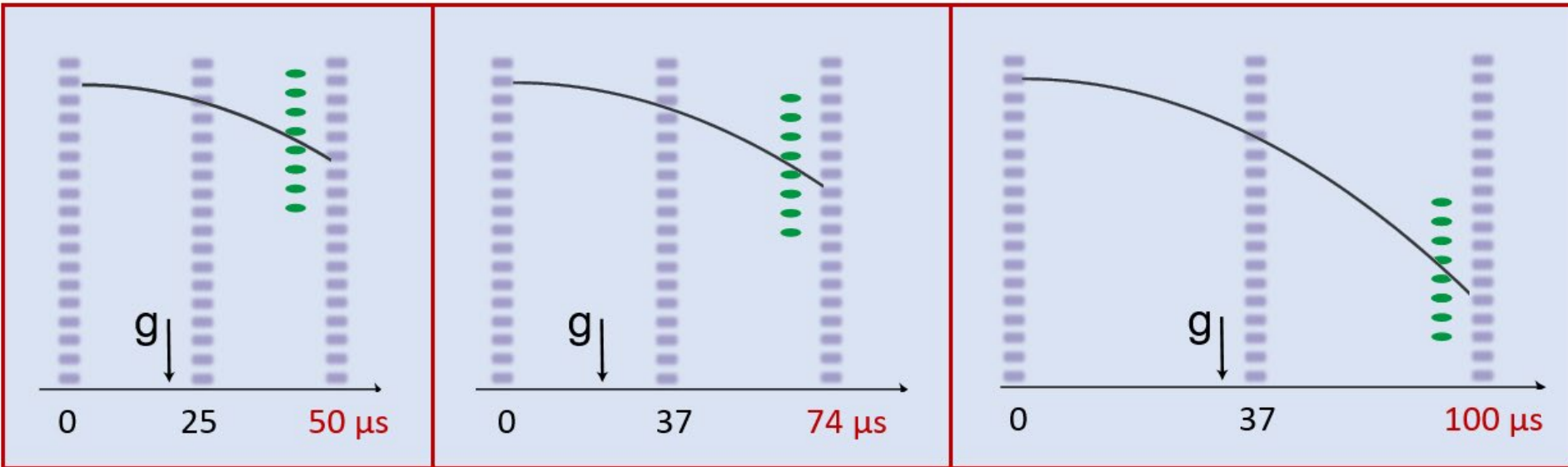


A. Shayeghi

P. Rieser

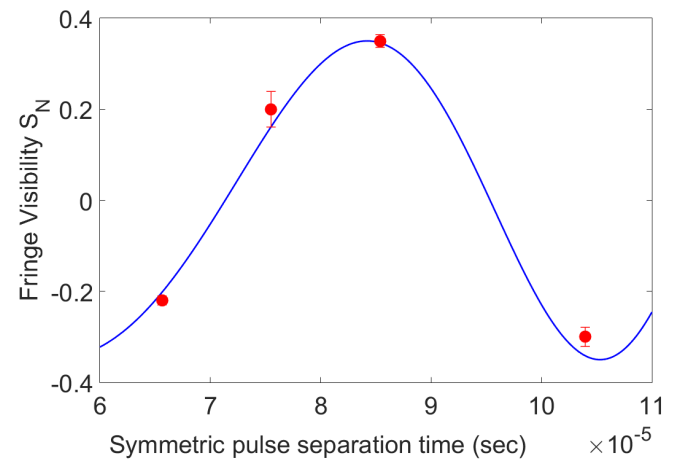


3. Molecular interference patterns as nanorulers to **measure free fall**: The OTIMA version of „COW“ interferometry

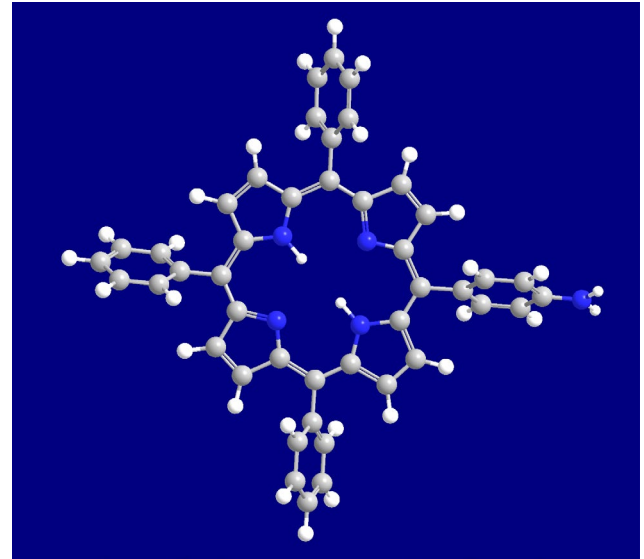
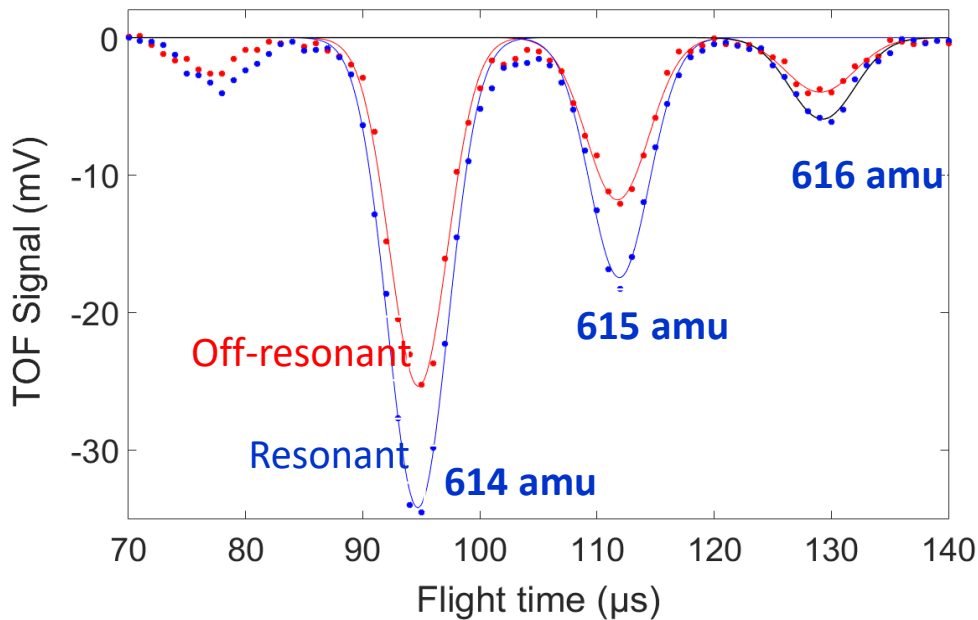


Symmetrically stretching the interferometer in time:

→ Gravitational acceleration “g”



Isotope-selective interference in free fall: Comparison of tetraphenylporphyrin isotopes



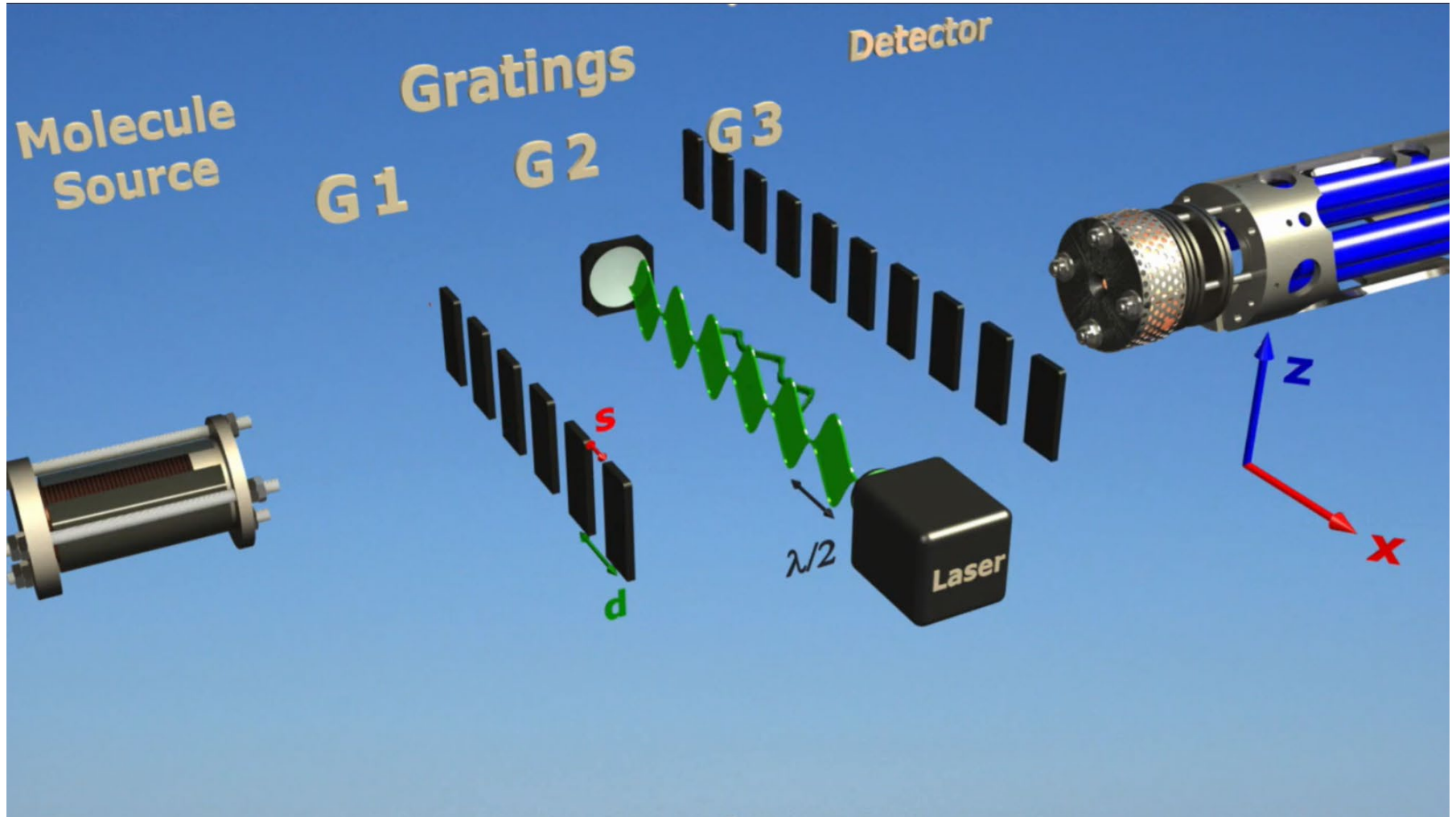
- **Isotopically pure bosonic TPP:** >50%
- Fermion with **one ^{13}C :** 30 %
- Boson with **two ^{13}C atoms:** 7%

- **Proof of principle:** $\eta := \Delta g/g < 1 \%$
- Future $T_{\text{flight}} = 1 \text{ s} \rightarrow \eta < 10^{-8}$
- **What can we learn from**
 - Ferromagnets & Superconductors?
 - High internal ring currents?
 - High internal energy?

**Long Base Line
Universal Matter-Wave
Interferometry**

LUMI 1.0 & upgrades

Long-Baseline Universal Matter-Wave Interferometry : **KDTLI / LUMI 1.0**



First $L = 0.2$ m interferometer: Gerlich et al. **Nature Physics** 3, 711 (2007)

First $L = 2.0$ m interferometer: Y. Y. Fein, P. Geyer, P. Zwick, F. Kiałka, S. Pedalino, M. Mayor, S. Gerlich & M. Arndt, **Nature Physics** 15, 1242 (2019).

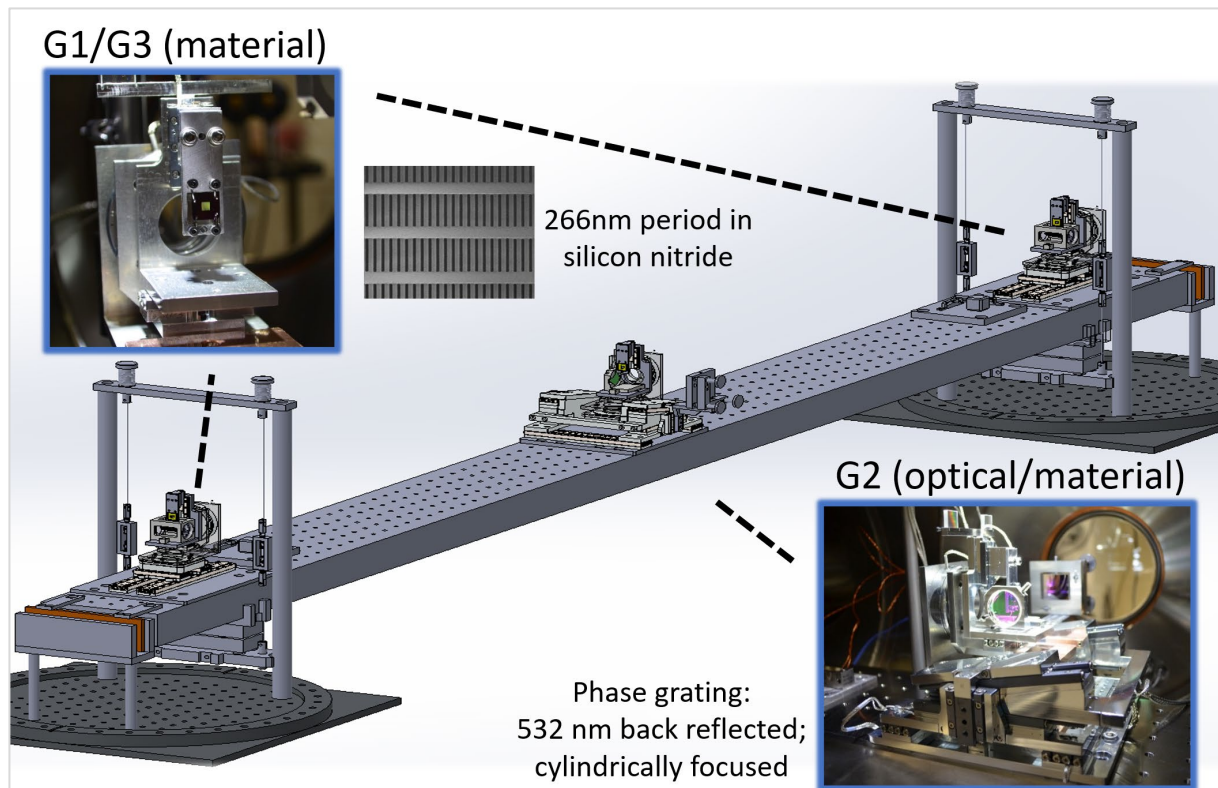
Long Base-line Universal Matter-Wave Interferometer: **LUMI**

1 - 100th order interference observed with:

- **Atoms:** Cs, Ba, Sr
- **Molecules:** C₆₀, Anthracene, Rubrene, ...
- **Tripeptides:** e.g. Ala-Trp-Ala
- **Macromolecules:** >25 kDa

Design specifications:

- **Wavelength** : $\lambda_{dB} < 50$ fm
- **Mass (@ 300 m/s):** $M > 3 \times 10^4$ amu
- **External force** : $F < 10^{-26}$ N



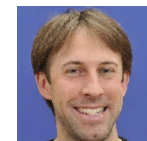
Y. Y. Fein



P. Geyer

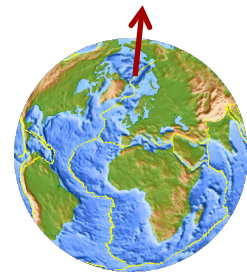


S. Pedalino



S. Gerlich

4. Fringe shifts caused by gravity and Earth's rotation



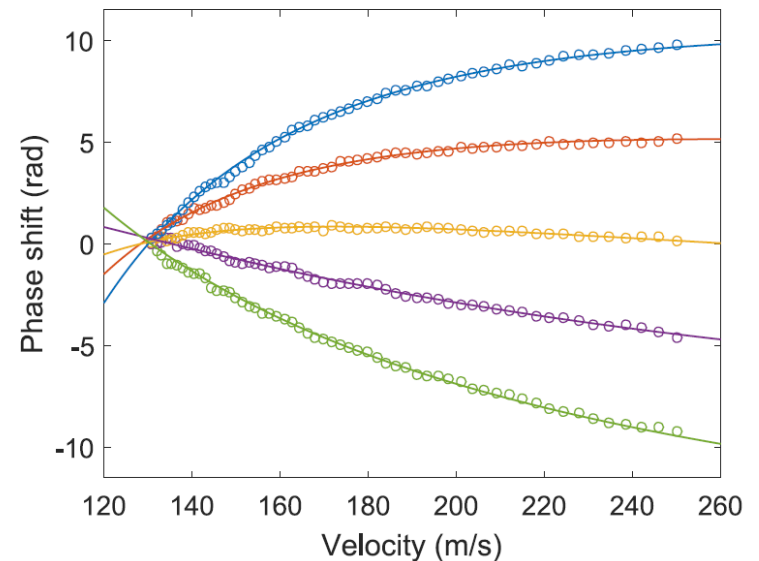
- In Talbot-Lau interferometry, the **acceleration** a yields a fringe phase shift $\Delta\phi = k_a a T^2$
- The flight time between the gratings is **velocity dependent**: $T = L/v$
- Averaging over a class of velocities reduces the interference contrast

$$\text{Earth's rotation: } \mathbf{a} = 2 \mathbf{v} \times \boldsymbol{\Omega}$$

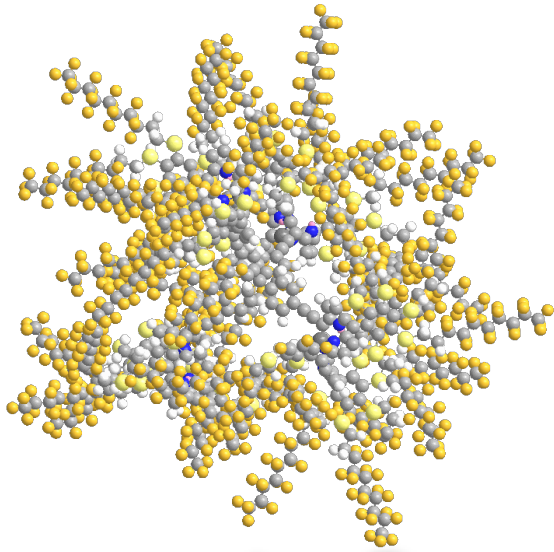
$$\text{Earth's gravity: } \mathbf{a} = \mathbf{g} \cdot \vartheta_{roll}$$

One can **compensate Coriolis & Gravity** acceleration by rolling the gratings by the angle:

$$\vartheta_{roll} = -\Omega_E v_p / g$$

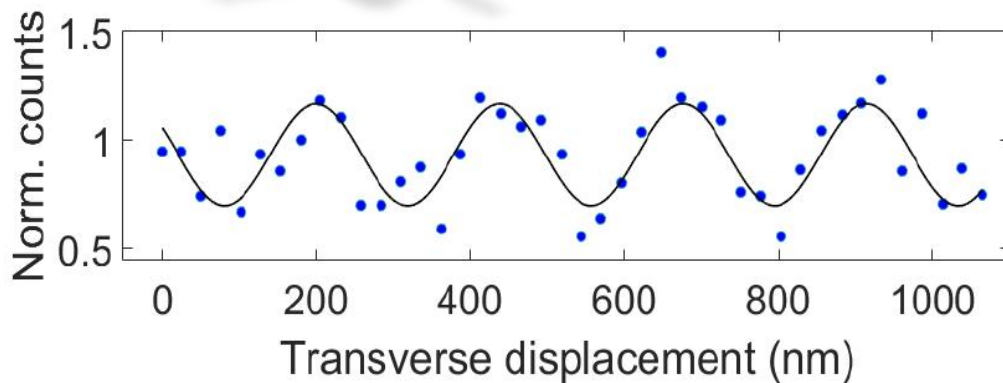


Quantum Wave Nature of the most massive molecular family



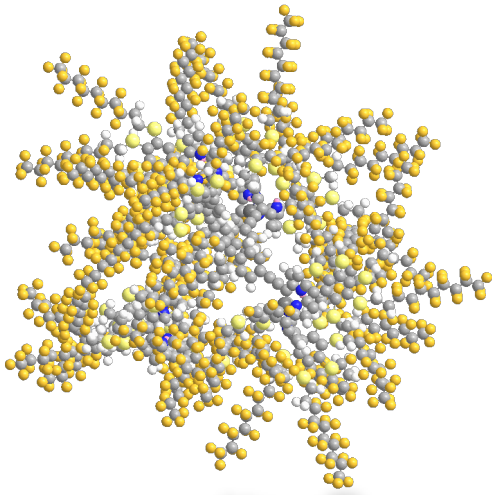
Perfluoroalkyl-functionalized Oligo-Porphyrines

- $m \simeq 25000 - 28000$ amu
- $N \simeq 1800 - 2000$ Atoms per molecule
- $C_{715}H_{260}F_{908}N_{16}S_{53}Zn_4$
- Velocity: $v = 300$ m/s
- Molecular diameter: $D = 50$ Å
- Billions of structural Isomers / Conformers
- De Broglie Wavelength $\lambda_{dB} = 50 - 60$ fm

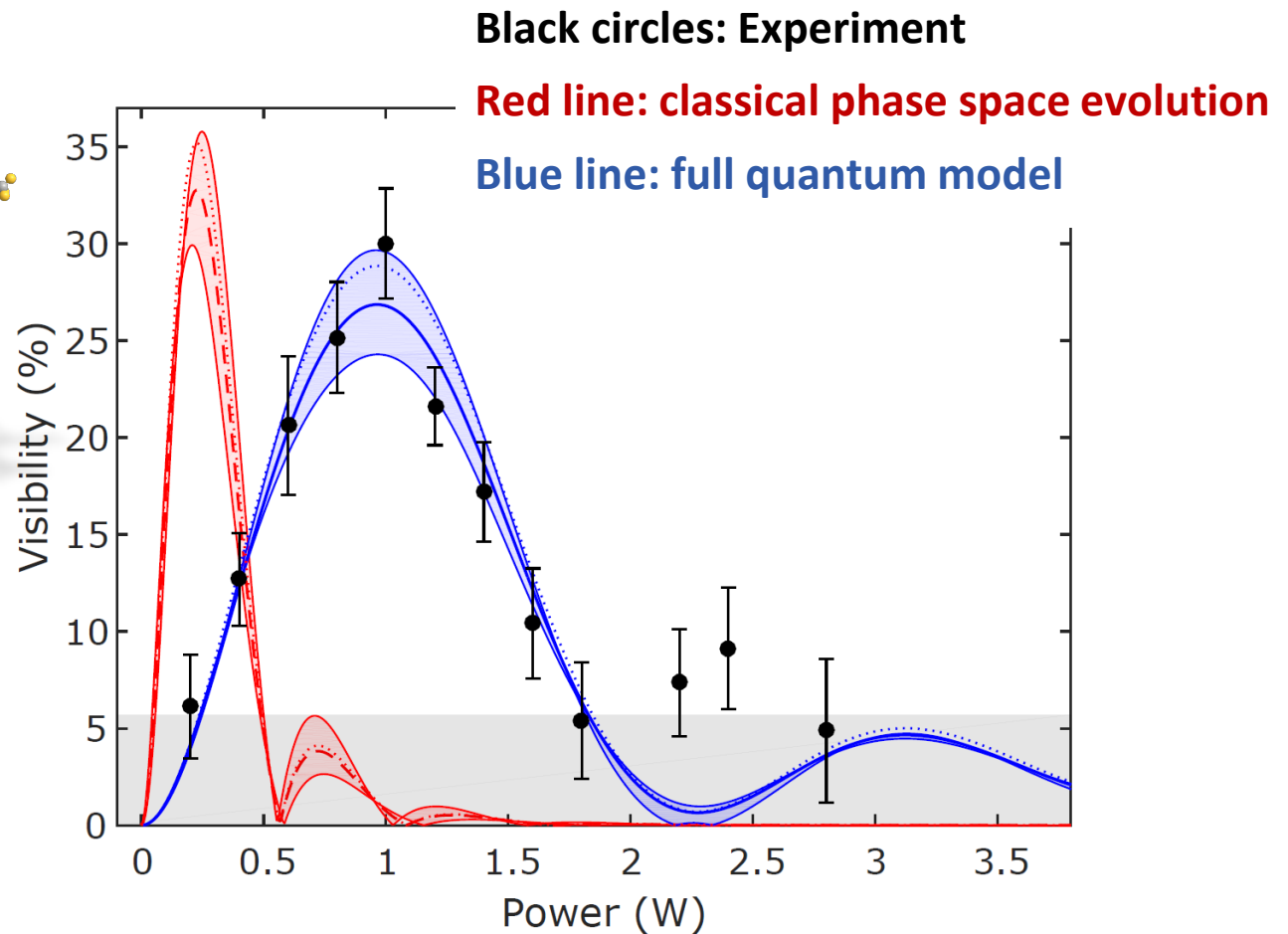


The interference pattern emerging when G3 is moved.

Matter-wave fringe visibility allows distinguishing Classical from Quantum Phenomena



- $m > 25000$ amu
- $N \approx 2000$ atoms

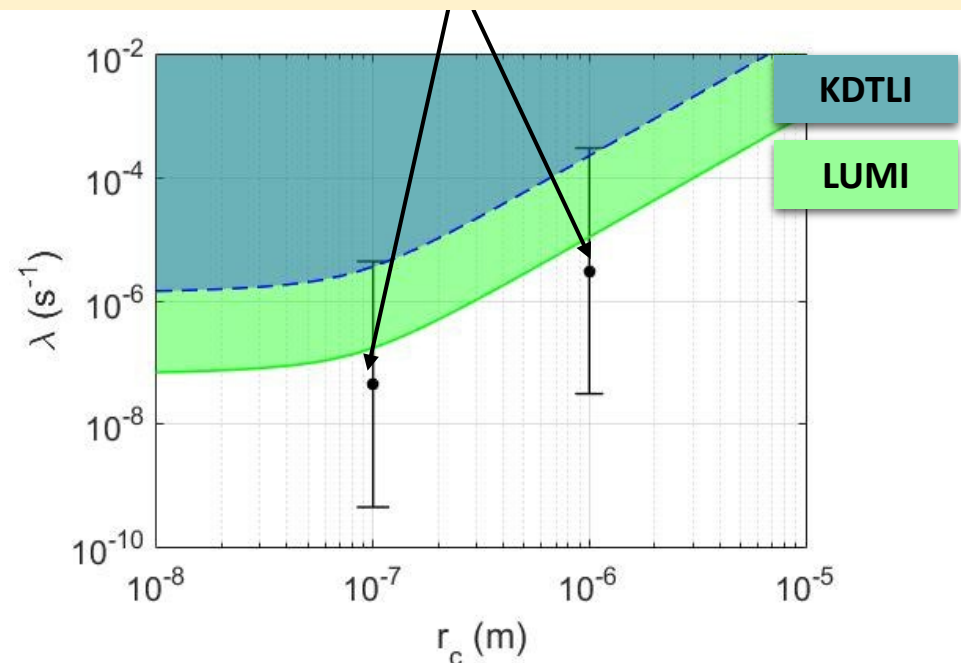


Interferometric bounds on Continuous Spontaneous Localization (CSL)

- CSL models assume a stochastic perturbation (by an unknown agent) which collapses the wave function at rate λ_c to a Gaussian state of width r_c .
- The collapse rate is predicted to scale like: $\lambda_c \propto m^2$
- If collapse occurred more often than expected, coherence would be lost.

LUMI sets currently the **strongest** interferometric bound on collapse models for $r_c \in [10 - 1000] \text{ nm}$

Prediction by S. Adler, J. Phys. A, 40 2935 (2007)



5. Quantum Macroscopicity

by S. Nimmrichter & K. Hornberger Phys. Rev. Lett. 110, 160403 (2013)

Quantum Macroscopicity $\mu \propto m^2$

measures how well
an experiment can **exclude**
a generic set of
Non-linear Extensions to
Quantum Mechanics

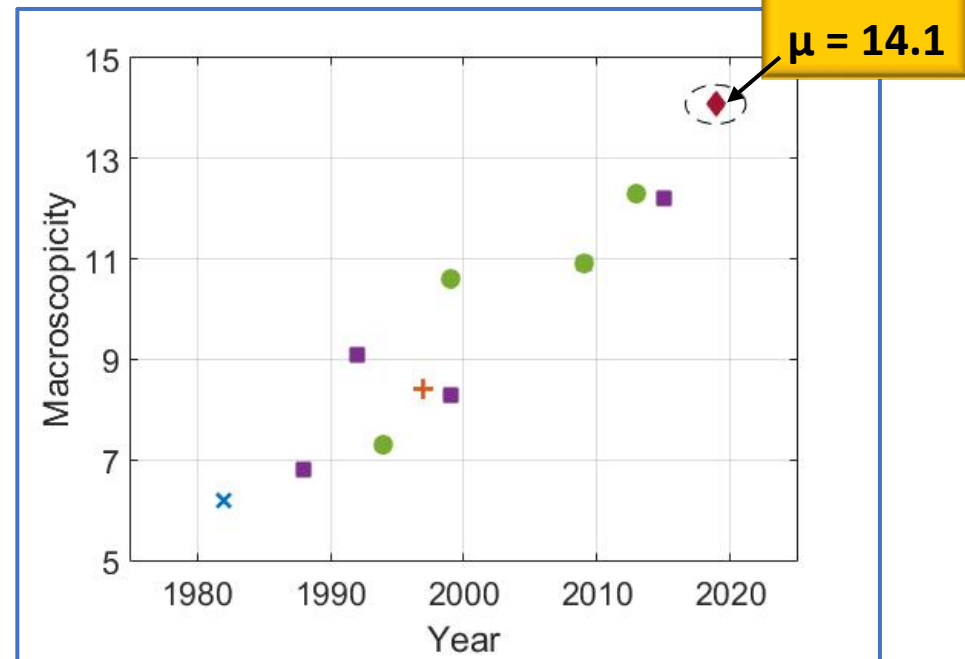
What does $\mu=14$ signify ?

To achieve the same bounds on QM
an **electron interferometer** would
need to maintain coherence (at
equal contrast and integration time)
for over **80 million years !**

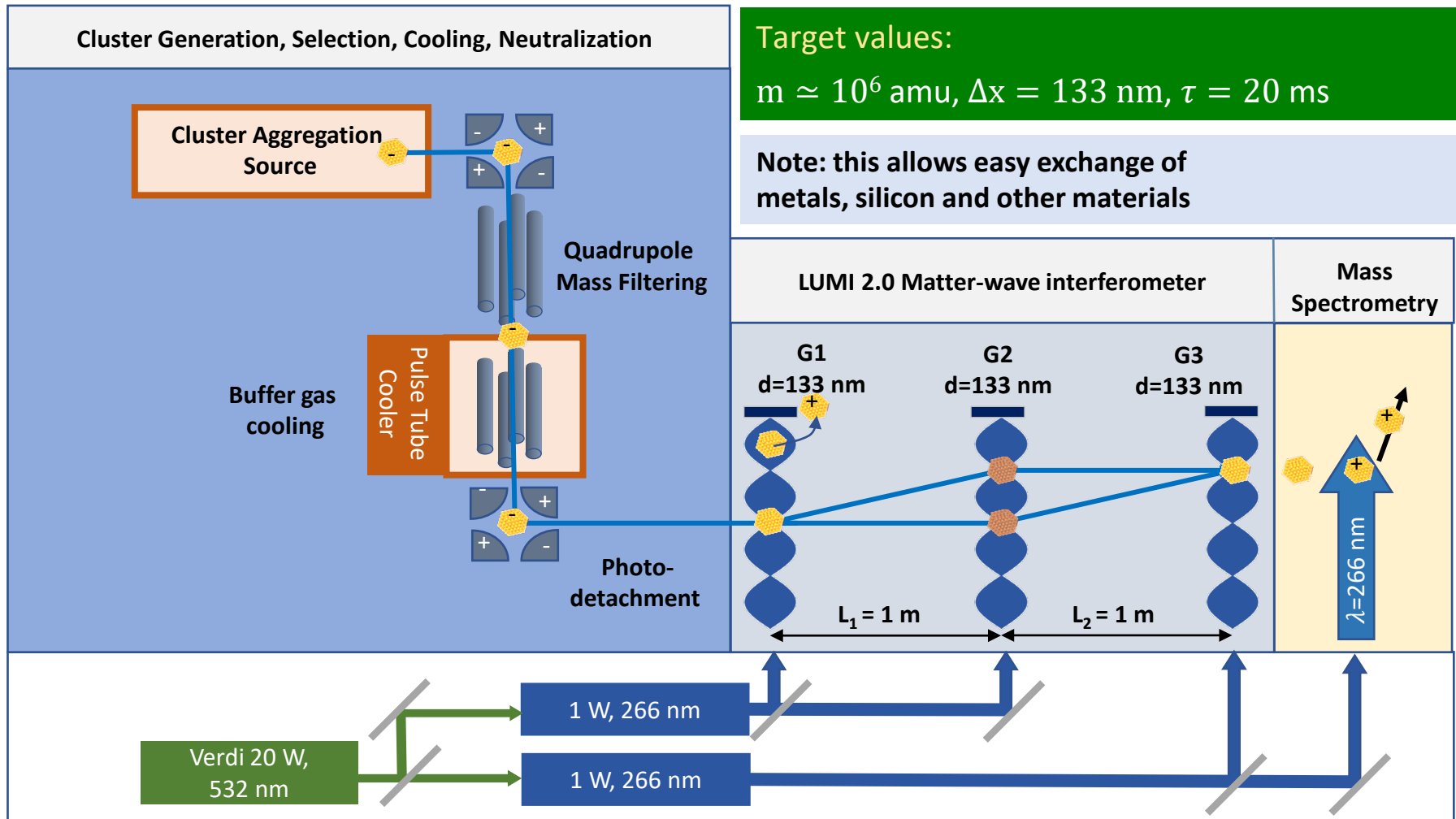
Previous macroscopicities for

- **Neutrons** : x
- **Atom** : ■
- **Molecules** : ●

$$\mu = \log_{10} \left(\frac{1}{\ln(V/V_0)} \left(\frac{m}{m_e} \right)^2 \frac{\tau}{1 \text{ sec}} \right)$$



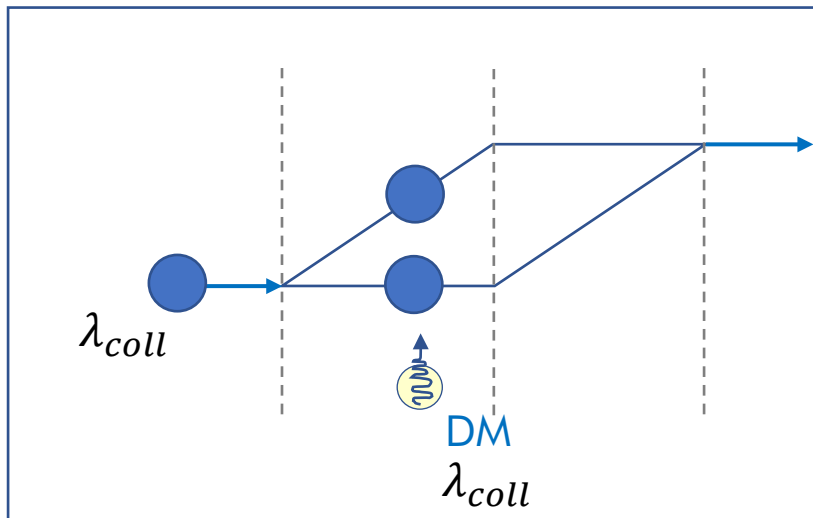
6. Work in progress: Interferometry with Large Metal Clusters: LUMI 2.0



Proposed upgrade to: $m \approx 10^8$ amu, $\Delta x = 80$ nm, $\tau = 2$ s in a tower on Earth

7. Low-energy dark matter: Why matter-waves may be favourable

- An interferometer with widely separated arms is sensitive to **decoherence** by **small momentum transfer** in collision with dark matter.
- Prepare widely separated quantum states (**Schrödinger Cat, Matter-Wave**)
If λ_{dB} (DM) < **wave-packet separation of P1** → „which-path information“
- If **coherent scattering**, d.h transverse coherence of DM > colliding matter particle:
collision cross section $\sigma_{coll} \propto m^2$... increases with mass.



Watch out for **daily** and **annually varying decoherence** caused by the motion of the Earth and the sun.

LUMI 3.0 target at QNP @ UNIVIE

$$M = 5 \times 10^5 \text{ amu}, \tau > 0.02 \text{ s}, \Delta V/V < 1\%$$

LUMI 4.0 perspectives for QNP@UNIVIE

$$M = 5 \times 10^7 \text{ amu}, \tau > 1 \text{ s}, \Delta V/V < 1\%$$

Towards high-mass quantum states

Cold source &

large coherent momentum transfer

8. Inelastic Kapitza-Dirac scattering & KD-blockade

The goal is to prepare ...

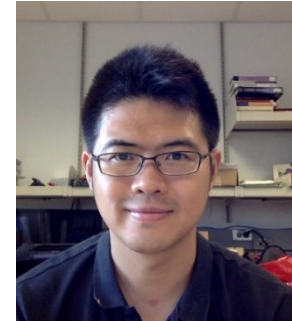
- Very large **Schrödinger Cat** states of massive particles
- Very **Large Momentum Transfer** Beam Splitters
- Nano-mechanical **number states**
- Nano-mechanical **2-level systems**

The method shall be universal and applicable to

- Trapped electrons
- Trapped atoms, trapped macromolecules & nanoparticles

Solution

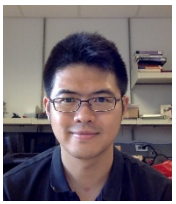
- Couple light & momentum and avoid internal states
Inelastic Kapitza Dirac scattering
- Suppress undesired transitions by amplitude interference
Kapitza-Dirac blockade



Wayne Cheng-Wei Huang
(Univ. of Göttingen)



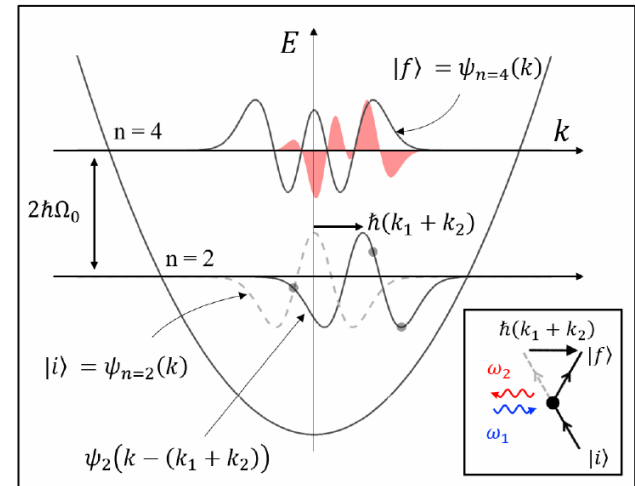
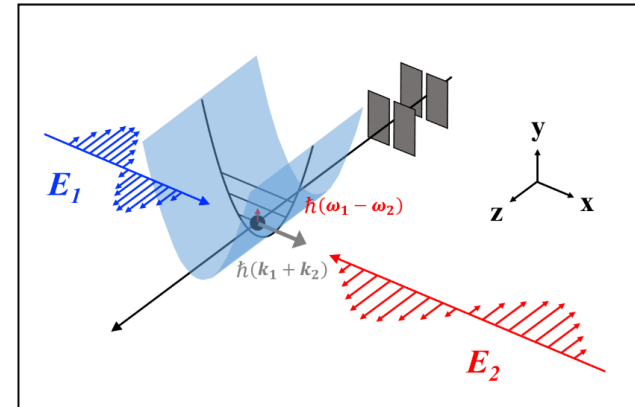
Herman Batelaan
(Lincoln, Nebraska)



Wayne Huang

Kapitza-Dirac blockade: A universal tool for selective state preparation

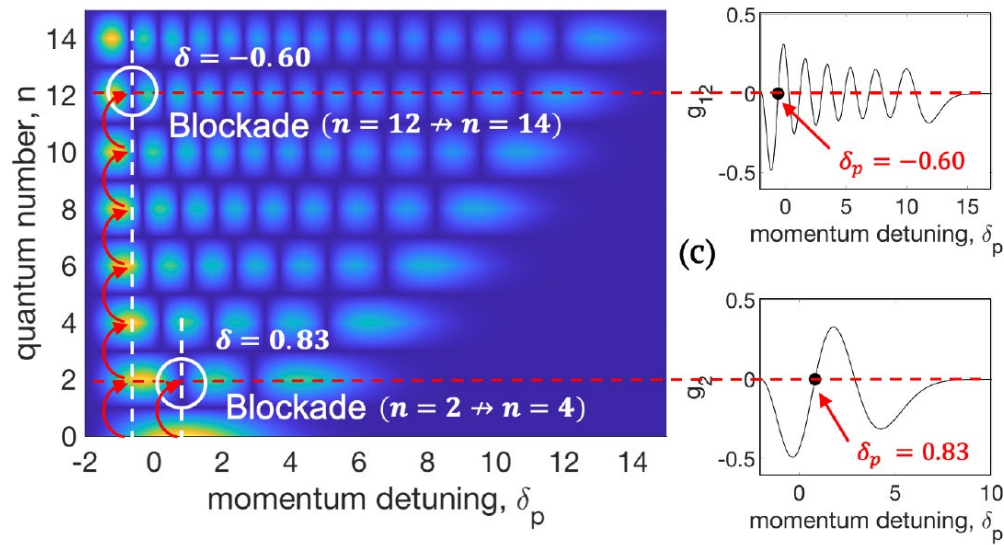
- **Massive particle in 1D trap:**
 - Optical tweezer
 - Electrodynamical trap ...
- **Population of ground state:**
 - Collimated beam
 - Ground state cooling (see Aspelmeyer talk)
- **Coherent state transfer by Kapitza-Dirac scattering:**
 - 2 counter-propagating laser beams:
 - Equal polarization, far-off resonance
 - Detuned by twice the harmonic oscillator frequency
 - Acting on
 - Charged Particles: **Ponderomotive potential**
 - Polarizable particles: **Dipole potential**
- **Kapitza-Dirac Blockade:**
 - Energy and momentum conservation
 - **Destructive Interference of scattering amplitudes** confines state transfer to the desired levels



Kapitza-Dirac Blockade: Transition amplitudes $g_n(\eta)$ vanish at certain values of the Lamb Dicke parameter (momentum detuning)

$$g_n(\eta) \equiv \langle n+2 | \cos((k_1 + k_2)x) | n \rangle = -\sqrt{\frac{n!}{(n+2)!}} \eta^2 L_n^{(2)}(\eta^2) e^{-\eta^2/2}$$

$$\eta \equiv (k_1 + k_2)x_0$$

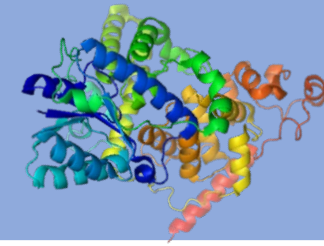


- Momentum splitting up to $\Delta p \approx 1000 \hbar k$ for electron & nanoparticle interferometry
- Rapid expansion of the wave function: **macroscopic cat of massive particles**
- $\Delta x \approx 400 \text{ nm}$ in 30 ms for $m=10^6 \text{ amu}$
- Can be tested with **macromolecule interferometry**

Summary & Outlook in to the Prospects of Universal matter-wave interferometry

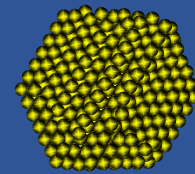
Large Biomolecules: 1 kDa - 10 MDa

Interface to Chemistry & Biology,
Quantum-enhanced metrology



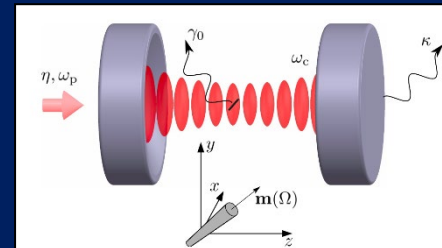
Large Metal Clusters: 0.1 - 100 MDa

Wave function Collapse, Decoherence,
Search for light dark matter



Dielectric Nanoparticles: 1 - 100 MDa

Rotational Cooling &
KD Blockade for non-trivial quantum states



Thank you for your
Attention 😊