

Proposal for testing the Schrödinger-Newton equation

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Institute for Theoretical Physics

Sao Paulo State University

In collaboration with

Gabriel Aguiar

Scientific Motivation

Low-energy quantum gravity

GENERAL RELATIVITY



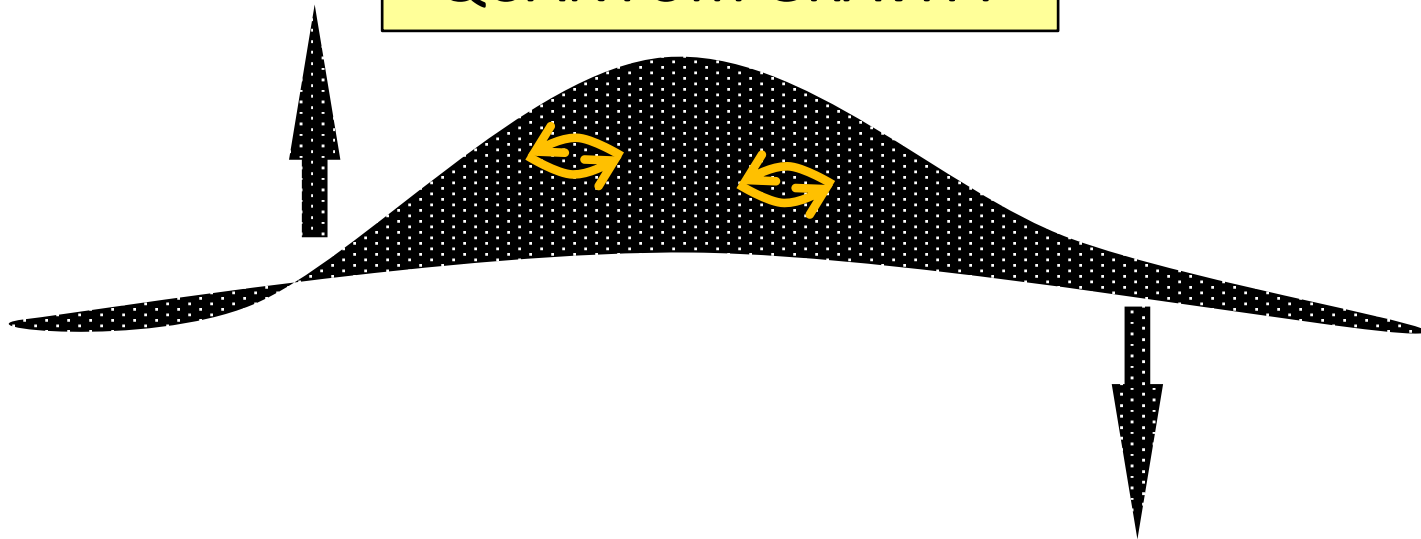
Spacetime and fields harmoniously evolved with classical rules

$$G_{ab}(g) = 8\pi T_{ab}(g, \Phi)$$

$$\nabla_a \nabla^a \Phi + m^2 \Phi = 0$$

Low-energy quantum gravity

QUANTUM GRAVITY



Spacetime and fields harmoniously evolved with quantum rules

?

Low-energy quantum gravity

QUANTUM FIELD THEORY



Fields evolved in a predefined spacetime with **quantum rules**

$$G_{ab}(g) = 8\pi T_{ab}(g, \Phi)$$

$$\nabla_a \nabla^a \Phi + m^2 \Phi = 0$$

$$\nabla_a \nabla^a \hat{\phi} = 0$$

Low-energy quantum gravity

SEMI-CLASSICAL GRAVITY



Backreaction of quantum fields in a classical spacetime

$$G_{ab}(g) = 8\pi T_{ab}(g, \Phi) + 8\pi \langle \hat{T}_{ab}(g, \hat{\phi}) \rangle$$

$$\nabla_a \nabla^a \Phi + m^2 \Phi = 0$$

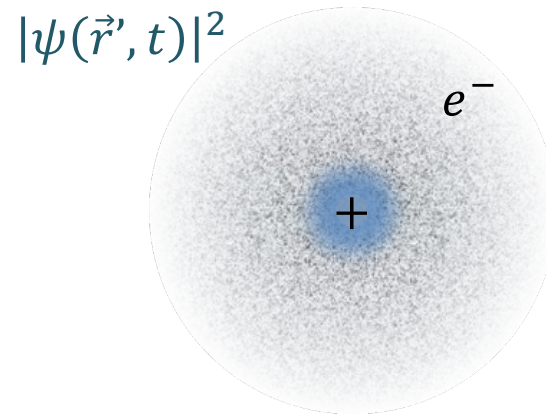
$$\nabla_a \nabla^a \hat{\phi} = 0$$

Schrödinger-Newton equation

Bagrami, Großardt, Donaldi, Bassi, NJP (2014)

$$i\hbar \frac{\partial}{\partial t} \psi(\vec{r}, t) = -\frac{\hbar^2}{2\mu} \nabla^2 \psi(\vec{r}, t) + V_{\text{ext}}(\vec{r}, t) \psi(\vec{r}, t) + U_{\text{self}}^{\text{grav}}(\vec{r}, t) \psi(\vec{r}, t)$$

$$U_{\text{self}}^{\text{grav}}(\vec{r}, t) = -Gm^2 \int d^3\vec{r}' \frac{|\psi(\vec{r}', t)|^2}{|\vec{r} - \vec{r}'|}$$



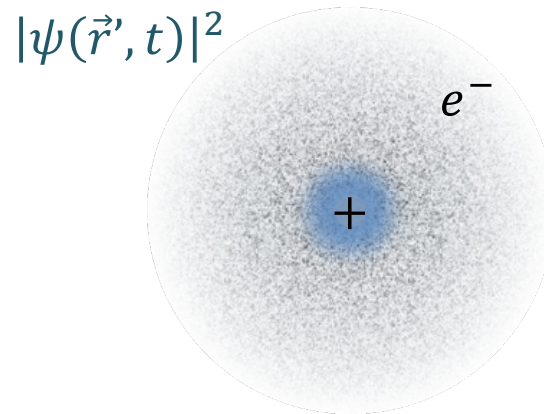
one-electron state

Schrödinger-Coulomb equation

da Silva, Aguiar, GEAM, PRA (2023)

$$i\hbar \frac{\partial}{\partial t} \psi(\vec{r}, t) = -\frac{\hbar^2}{2\mu} \nabla^2 \psi(\vec{r}, t) + V_{\text{ext}}(\vec{r}, t) \psi(\vec{r}, t) + U_{\text{self}}^{\text{elec}}(\vec{r}, t) \psi(\vec{r}, t)$$

$$U_{\text{self}}^{\text{elec}}(\vec{r}, t) = k_e e^2 \int d^3\vec{r}' \frac{|\psi(\vec{r}', t)|^2}{|\vec{r} - \vec{r}'|}$$



one-electron state

LEVEL	Schrödinger equation	Schrödinger-Coulomb equation
n=1	-13.598 eV	-1.2561 eV

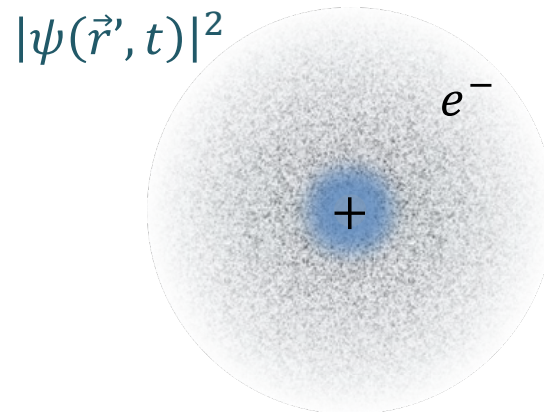
Schrödinger-Newton equation

da Silva, Aguiar, GEAM, PRA (2023)

$$i\hbar \frac{\partial}{\partial t} \psi(\vec{r}, t) = -\frac{\hbar^2}{2\mu} \nabla^2 \psi(\vec{r}, t) + V_{\text{ext}}(\vec{r}, t) \psi(\vec{r}, t) + U_{\text{self}}^{\text{grav}}(\vec{r}, t) \psi(\vec{r}, t)$$

$$U_{\text{self}}^{\text{grav}}(\vec{r}, t) = -Gm^2 \int d^3\vec{r}' \frac{|\psi(\vec{r}', t)|^2}{|\vec{r} - \vec{r}'|}$$

$$\frac{Gm_e^2}{k_e e^2} \sim 10^{-43}$$



one-electron state

Awfully difficult to test the Schrödinger-Newton equation with stationary states

Probing the Schrödinger-Newton equation in a Stern-Gerlach like experiment

Aguiar, GEAM, PRA (2024)

see also Großardt, PRD (2024)

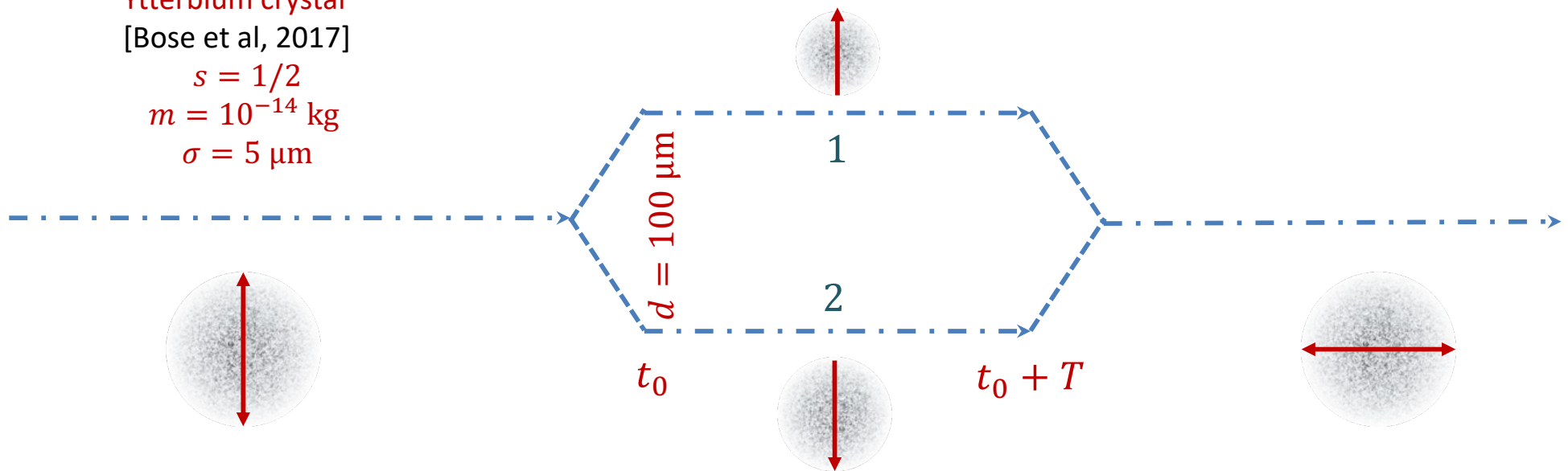
Schrödinger-Newton equation

$$i\hbar \frac{\partial}{\partial t} \psi_{\uparrow(\downarrow)}(t) \approx U_{\text{self}}^{\text{grav}}(\vec{r}_{\uparrow(\downarrow)}, t_0) \psi_{\uparrow(\downarrow)}(t)$$

$$U_{\text{self}}^{\text{grav}}(\vec{r}_{\uparrow(\downarrow)}, t) = -Gm^2 \int d^3\vec{r}' \frac{|\psi_{\uparrow(\downarrow)}(\vec{r}', t)|^2}{|\vec{r}_{\uparrow(\downarrow)} - \vec{r}'|}$$

One-particle state
Ytterbium crystal
[Bose et al, 2017]

$s = 1/2$
 $m = 10^{-14}$ kg
 $\sigma = 5 \mu\text{m}$

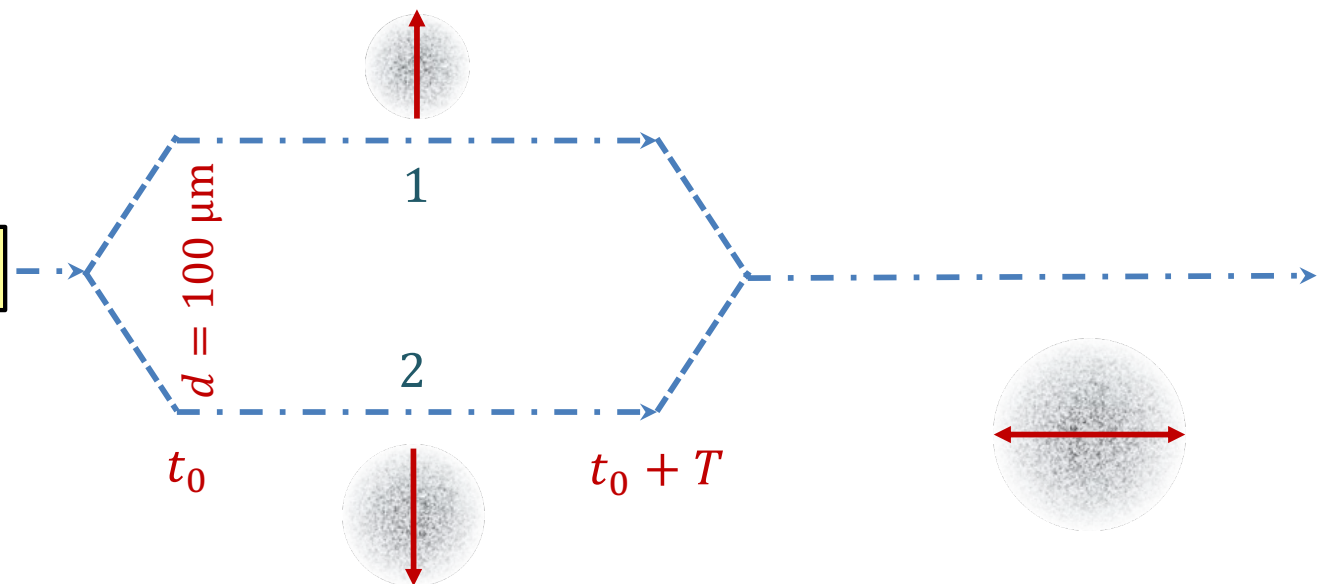
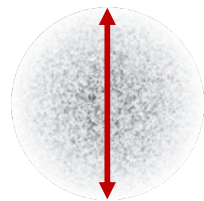


Schrödinger-Newton equation

Aguiar, GEAM, PRA (2024)

One-particle state
Ytterbium crystal
[Bose et al, 2017]
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$$G_\sigma(\vec{r}) (\cos\beta |\uparrow\rangle + \sin\beta |\downarrow\rangle)$$



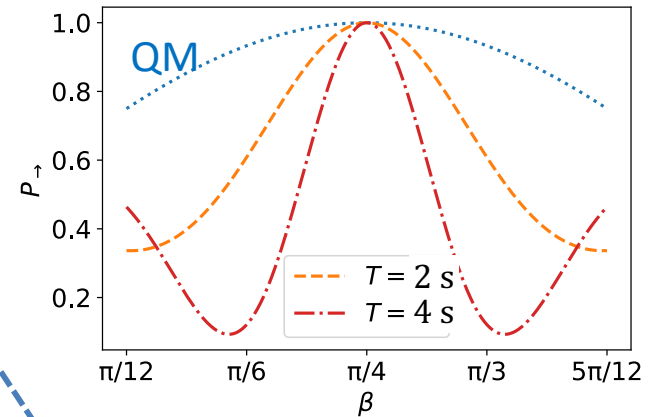
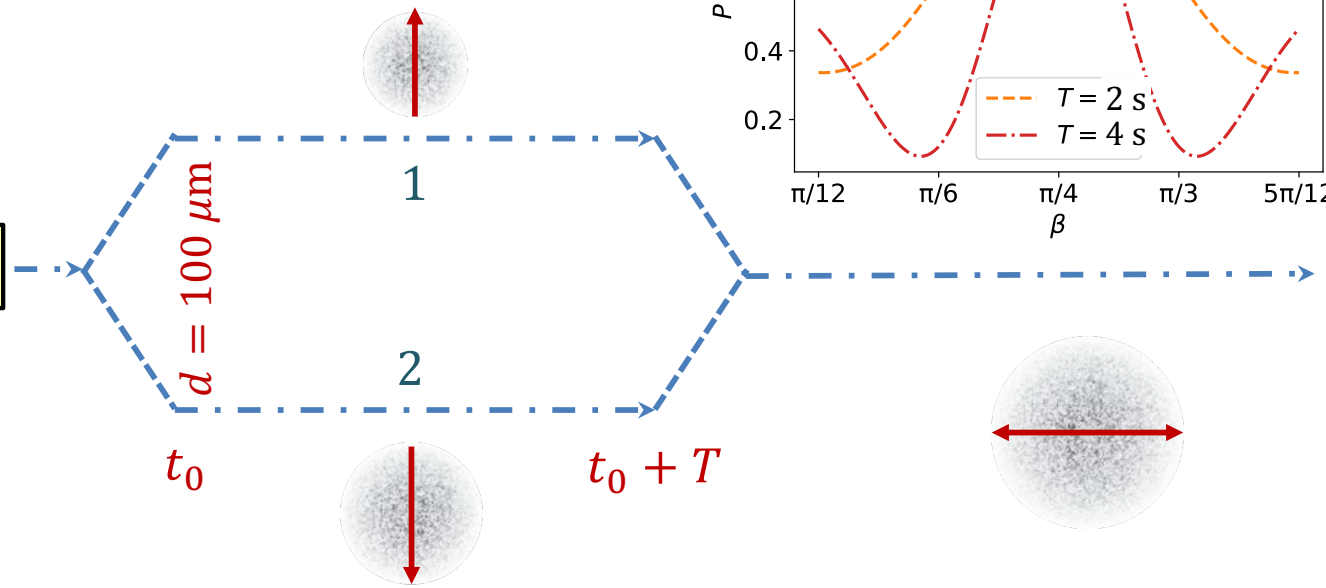
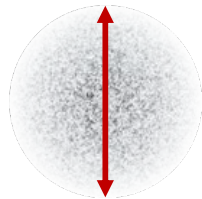
Schrödinger-Newton equation

Aguiar, GEAM, PRA (2024)

$$\Delta\phi = 2 \times 10^{14} \left(\frac{m}{m_p}\right)^2 \frac{(T/1s)}{(\sigma/1\mu m)}$$

One-particle state
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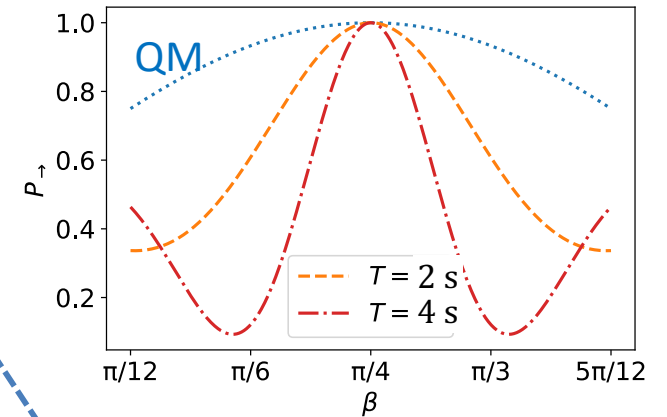
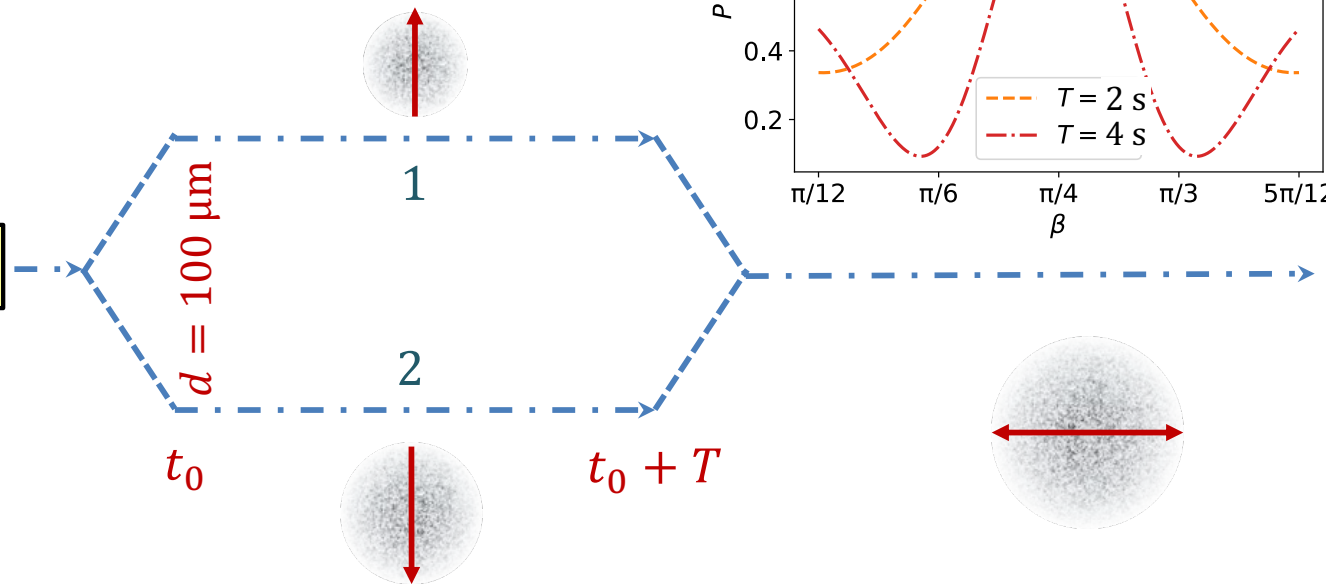
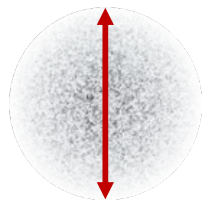
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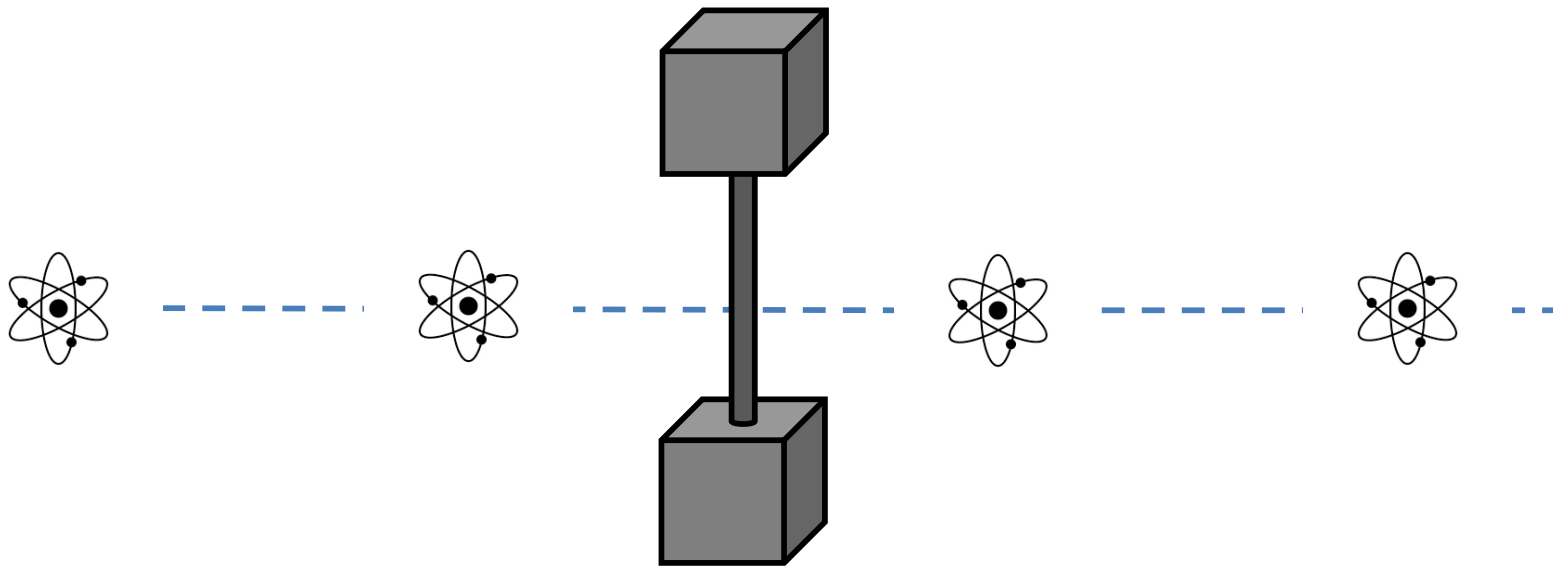
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Features of Semiclassical Gravity

GENERAL RELATIVITY

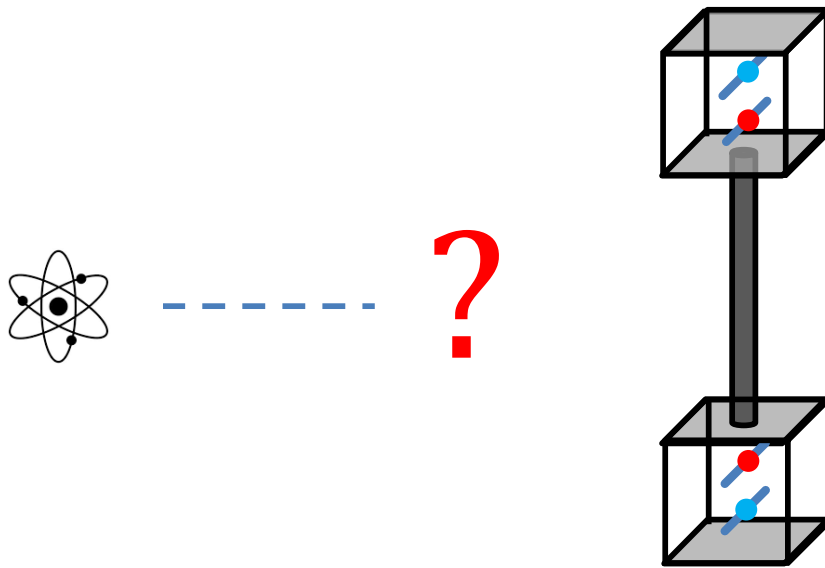


$$G_{ab}(g) = 8\pi T_{ab}(g, \Phi)$$

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Features of Semiclassical Gravity

QUANTUM GRAVITY

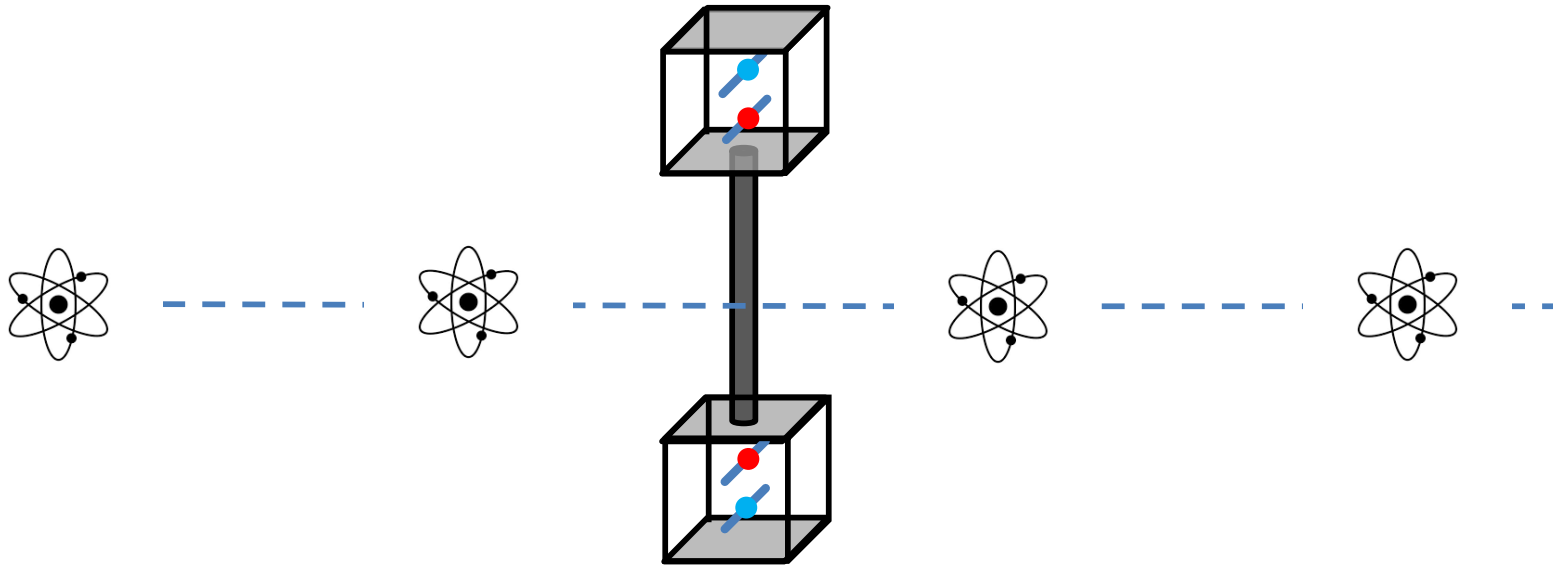


$$|\psi\rangle = \frac{1}{\sqrt{2}} (|E_0, E\rangle + |E, E_0\rangle)$$

$$\langle E_0 | E_0 \rangle = \langle E | E \rangle = 1, \quad \langle E_0 | E \rangle = 0$$

Features of Semiclassical Gravity

SEMI-CLASSICAL GRAVITY

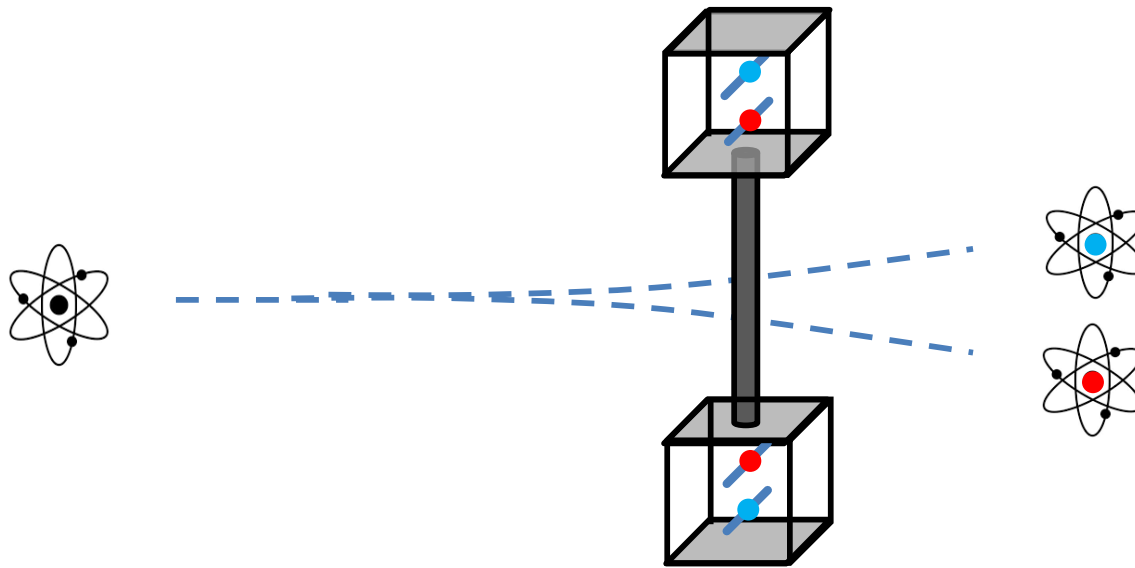


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Features of Semiclassical Gravity

QUANTUM GRAVITY



$$|\psi\rangle = \frac{1}{\sqrt{2}} (|E_0, E, \text{red nucleus}\rangle + |E, E_0, \text{blue nucleus}\rangle)$$

$$\langle E_0 | E_0 \rangle = \langle E | E \rangle = 1, \quad \langle E_0 | E \rangle = 0$$